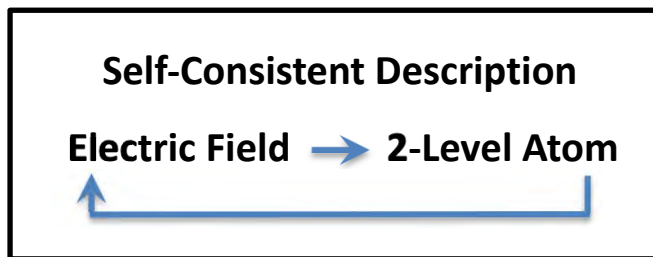


# Maxwell-Bloch Equations

So far in the Semiclassical Description

- (\*) Classical light acting on quantum atoms
- (\*) Next: Close the loop



We need to set up and solve a set of workable simultaneous equations for the atoms and field.

- (1) The electric field. We write

$$\vec{E}(z, t) = \vec{\epsilon} \mathcal{E}(z, t) e^{-i(\omega t - kz)}$$

↑  
wavepacket envelope

- Plane wave propagating in the  $+z$  direction, the real part is the physical field

Slowly Varying Envelope Approximation (SVEA)

↓

We require that the envelope  $\mathcal{E}(z, t)$  is smooth in space and time compared to the plane wave part.

↓

$$\begin{aligned} \left| \frac{\partial \mathcal{E}}{\partial z} \right| &\ll k |\mathcal{E}| & \left| \frac{\partial \mathcal{E}}{\partial t} \right| &\ll \omega |\mathcal{E}| \\ \left| \frac{\partial^2 \mathcal{E}}{\partial z^2} \right| &\ll k \left| \frac{\partial \mathcal{E}}{\partial z} \right| & \left| \frac{\partial^2 \mathcal{E}}{\partial z^2} \right| &\ll \omega \left| \frac{\partial \mathcal{E}}{\partial t} \right| \end{aligned}$$

This is not particularly restrictive, unless working with ultrafast lasers.

- (2) The Macroscopic Polarization Density.

We use the quantum expectation value  $\vec{P}(z, t) = N \langle \hat{p} \rangle$

Of this, we need the complex part that goes with  $\vec{E}(z, t)$  and can be plugged into the wave equation.

# Maxwell-Bloch Equations

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This is not particularly restrictive, unless working with ultrafast lasers.

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We use the quantum expectation value  $\vec{P}(z,t) = N \langle \hat{\vec{p}} \rangle$

Of this, we need the complex part that goes with  $\vec{\mathcal{E}}(z,t)$  and can be plugged into the wave equation.

Thus, of

$$\begin{aligned} \langle \hat{\vec{p}} \rangle &= \vec{\eta}_{12} \langle a_2 a_1^\dagger \rangle + \vec{\eta}_{21} \langle a_1 a_2^\dagger \rangle \\ &= \vec{\eta}_{12} \mathcal{S}_{21} e^{-i(\omega t - kz)} + \vec{\eta}_{21} \mathcal{S}_{12} e^{i(\omega t - kz)} \end{aligned}$$

← slow variables

we need the part that goes as  $e^{-i(\omega t - kz)}$

The *physical field* is  $\text{Re}[\vec{\mathcal{E}} \mathcal{E}(z,t) e^{-i(\omega t - kz)}]$

The *physical dipole* is

$$\text{Re}[\vec{\eta}_{12} \mathcal{S}_{21} e^{-i(\omega t - kz)} + \vec{\eta}_{21} \mathcal{S}_{12} e^{i(\omega t - kz)}]$$

Note factor of 2  $\rightarrow = 2 \times \text{Re}[\vec{\eta}_{12} \mathcal{S}_{21} e^{-i(\omega t - kz)}]$

Note: The coherence  $\mathcal{S}_{21}$  depends on  $z,t$  because the field depends on  $z,t$  through the envelope  $\mathcal{E}(z,t)$   $\rightarrow$  implicit SVEA on  $\mathcal{S}_{21}$ .

Note: In a real, multilevel atom  $\vec{\eta}_{12}$  need not be parallel to the field. However, only the part that is parallel to the field can emit radiation that interferes with it and lead to absorption and dispersion.

# Maxwell-Bloch Equations

Thus, of

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The complex dipole parallel to  $\vec{E}$  is

$$\begin{aligned}\vec{P}(z, t) &= \vec{E} 2N \underbrace{(\vec{\mu}_{12} \cdot \vec{E}^*)}_{\mu^*} \mathcal{S}_{21}(z, t) e^{-i(\omega t - k z)} \\ &\equiv \vec{E} 2N \mu^* \mathcal{S}_{21}(z, t) e^{-i(\omega t - k z)}\end{aligned}$$

**Final Note:** Because of the RWA we have

$$\left| \frac{\partial \mathcal{S}_{21}}{\partial t} \right| \ll \omega |\mathcal{S}_{21}| \quad \left| \frac{\partial^2 \mathcal{S}_{21}}{\partial t^2} \right| \ll \omega \left| \frac{\partial \mathcal{S}_{21}}{\partial t} \right|$$

**(3) Maxwells eqs.  $\Rightarrow$  Wave Equation**

$$\left( \frac{\partial^2}{\partial z^2} - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) \vec{E}(z, t) = \frac{1}{\epsilon_0 c^2} \frac{\partial^2}{\partial t^2} \vec{P}(z, t)$$

We plug in the complex  $\vec{E}(z, t)$  and  $\vec{P}(z, t)$ , use the SVEA conditions on the derivatives to eliminate all but the leading terms, and finally take the scalar product with  $\vec{E}$

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# Maxwell-Bloch Equations

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This gives us our final equation for the envelope:

$$\begin{aligned}\left( \frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t} \right) \mathcal{E}(z,t) &= \frac{ik}{\epsilon_0} N \mu^* \mathcal{S}_{21}(z,t) \\ \text{where } \mu^* &= \vec{\mu}_{12} \cdot \vec{\mathcal{E}}^*\end{aligned}$$

Write  $\mathcal{S}_{21}$  in terms of the Bloch variables to get the

## Maxwell-Bloch Equations

$$\begin{aligned}\left( \frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t} \right) \mathcal{E}(z,t) &= \frac{ik}{2\epsilon_0} N \mu^* (\mu - i\nu) \\ \dot{\mu} &= -\beta\mu + \text{Im}[\chi]\omega + \Delta\nu \\ \dot{\nu} &= -\beta\nu + \Delta\mu + \text{Re}[\chi]\omega \\ \dot{\omega} &= -\frac{1}{T_1} (\nu + \omega) - \text{Re}[\chi]\nu - \text{Im}[\chi]\mu\end{aligned}$$

**Note:** The Maxwell-Bloch Equations are a key result. They lead to rich physics, including absorption, gain, dispersion, solitons, lasers, and much more.

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## Maxwell-Bloch Equations

$$\left(\frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t}\right) \mathcal{E}(z, t) = \frac{ik}{2\epsilon_0} N \mu^* (u - i v)$$

$$\dot{u} = -\beta u + \text{Im}[\chi] \omega + \Delta v$$

$$\dot{v} = -\beta v + \Delta u + \text{Re}[\chi] \omega$$

$$\dot{\omega} = -\frac{1}{T_1} (\gamma + \omega) - \text{Re}[\chi] v - \text{Im}[\chi] u$$

**Note:** The Maxwell-Bloch Equations are a key result. They lead to rich physics, including absorption, gain, dispersion, self-induced transparency, solitons, lasers, and much more.

## Steady-State Solutions to MBE's

Steady state means that

$$\frac{1}{c} \frac{\partial \mathcal{E}}{\partial t} = 0 \quad \& \quad \mathcal{S}_{21}(z, t) \rightarrow \mathcal{S}_{21}(z, \infty) = \frac{-i\chi/2}{\beta + i\Delta} (\mathcal{S}_{22} - \mathcal{S}_{11})$$

Combine with  $\chi = -\vec{\mu}_{21} \cdot \vec{\mathcal{E}}/\hbar = \mu \mathcal{E}/\hbar$



$$\begin{aligned} \frac{\partial \mathcal{E}}{\partial z} &= \frac{ik}{\epsilon_0} N \mu^* \left(\frac{-i\mu \mathcal{E}}{2\hbar}\right) \frac{1}{\beta + i\Delta} (\mathcal{S}_{22} - \mathcal{S}_{11}) \\ &= \frac{\hbar N |\mu|^2}{2\hbar \epsilon_0} \frac{\beta - i\Delta}{\Delta^2 + \beta^2} (\mathcal{S}_{22} - \mathcal{S}_{11}) \mathcal{E} \end{aligned}$$

We can rewrite this as

$$\frac{\partial \mathcal{E}}{\partial z} = \frac{1}{2} (a - i\delta) \omega \mathcal{E}$$

$$a = \frac{\hbar N |\mu|^2}{2\hbar \epsilon_0} \frac{\beta}{\Delta^2 + \beta^2} = N \sigma(\Delta)$$

$$\delta = \frac{\hbar N |\mu|^2}{2\hbar \epsilon_0} \frac{\Delta}{\Delta^2 + \beta^2} = N \frac{\Delta}{\beta} \sigma(\Delta)$$

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To compare with our classical theory of dispersion, we solve for  $\mathcal{E}(z)$  and plug into eq. for a plane wave.

$$\begin{aligned} \text{Field:} \quad & E(z) = \mathcal{E}(z) e^{ikz} \\ \text{Envelope:} \quad & \mathcal{E}(z) = \mathcal{E}(0) e^{\left(\frac{a\omega}{2}\right)z} e^{i\left(-\frac{\delta\omega}{2}\right)z} \end{aligned} \quad \left. \vphantom{\begin{aligned} \text{Field:} \\ \text{Envelope:} \end{aligned}} \right\} \rightarrow$$

$$\text{Field:} \quad E(z) = \mathcal{E}(0) e^{\left(\frac{a\omega}{2}\right)z} e^{i\left(1 - \frac{\delta\omega}{2k}\right)kz}$$

$$\text{Compare to:} \quad E(z) = E_0 e^{-n_I k z} e^{in_R k z}$$

## Real & Imaginary Index of Refraction

$$\begin{aligned} n_I &= -\frac{a\omega}{2k} = -\frac{N\omega}{2k} \sigma(\Delta) \\ n_R &= 1 - \frac{\delta\omega}{2k} = 1 - \frac{\Delta}{\beta} \frac{N\omega}{2k} \sigma(\Delta) \end{aligned}$$

Analogous to results from Electron Oscillator

$$n_I(\omega) = \frac{Ne^2}{4\epsilon_0 m_e \omega} = \frac{\beta}{\Delta^2 + \beta^2}, \quad n_R(\omega) = 1 + \frac{\Delta}{\beta} n_I(\omega)$$

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## Behavior of the Intensity

$$\begin{aligned} \frac{\partial}{\partial z} |\mathcal{E}^* \mathcal{E}| &= \mathcal{E}^* \frac{\partial \mathcal{E}}{\partial z} + \frac{\partial \mathcal{E}^*}{\partial z} \mathcal{E} \\ &= \frac{1}{2} (a - id)\omega |\mathcal{E}|^2 + \frac{1}{2} (a + id)\omega |\mathcal{E}|^2 = a\omega |\mathcal{E}|^2 \end{aligned}$$



$$\frac{\partial I}{\partial z} = a\omega I = a(\mathcal{G}_{21} - \mathcal{G}_{11}) I$$

Note that  $\left\{ \begin{array}{l} a = N\sigma(\Delta) \geq 0 \\ I(z) = I(0) e^{a(\mathcal{G}_{21} - \mathcal{G}_{11})z} \end{array} \right.$



Exp. Decay of  $I$  for  $\mathcal{G}_{21} - \mathcal{G}_{11} < 0$

Exp. growth of  $I$  for  $\mathcal{G}_{21} - \mathcal{G}_{11} > 0$

must be maintained by some external process



# Maxwell-Bloch Equations

To compare with our classical theory of dispersion, we solve for  $\mathcal{E}(z)$  and plug into eq. for a plane wave.

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Compare to:  $E(z) = E_0 e^{-n_I k z} e^{i n_R k z}$



## Real & Imaginary Index of Refraction

$$n_I = -\frac{a\omega}{2k} = -\frac{N\omega}{2k} \sigma(\Delta)$$

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$$n_I(\omega) = \frac{Ne^2}{4\epsilon_0 m_e \omega} = \frac{\beta}{\Delta^2 + \beta^2}, \quad n_R(\omega) = 1 + \frac{\Delta}{\beta} n_I(\omega)$$

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$$\begin{aligned} \frac{\partial}{\partial z} |\mathcal{E}^* \mathcal{E}| &= \mathcal{E}^* \frac{\partial \mathcal{E}}{\partial z} + \frac{\partial \mathcal{E}^*}{\partial z} \mathcal{E} \\ &= \frac{1}{2} (a - id)\omega |\mathcal{E}|^2 + \frac{1}{2} (a + id)\omega |\mathcal{E}|^2 = a\omega |\mathcal{E}|^2 \end{aligned}$$



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# Maxwell-Bloch Equations

## Behavior of the Intensity

$$\frac{\partial}{\partial z} |\mathcal{E}^* \mathcal{E}| = \mathcal{E}^* \frac{\partial \mathcal{E}}{\partial z} + \frac{\partial \mathcal{E}^*}{\partial z} \mathcal{E}$$

$$= \frac{1}{2} (a - id) \omega |\mathcal{E}|^2 + \frac{1}{2} (a + id) \omega |\mathcal{E}|^2 = a \omega |\mathcal{E}|^2$$



$$\frac{\partial I}{\partial z} = a \omega I = a (\rho_{22} - \rho_{11}) I$$

Note that

$$\begin{cases} a = N \sigma(\Delta) \geq 0 \\ I(z) = I(0) e^{a(\rho_{22} - \rho_{11})z} \end{cases}$$



Exp. Decay of  $I$  for  $\rho_{22} - \rho_{11} < 0$   
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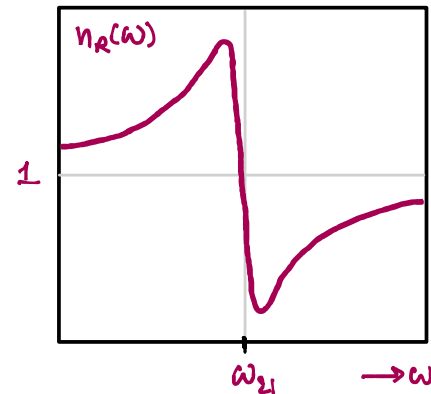
## Behavior of the Dispersion:

### Real & Imaginary Index of Refraction

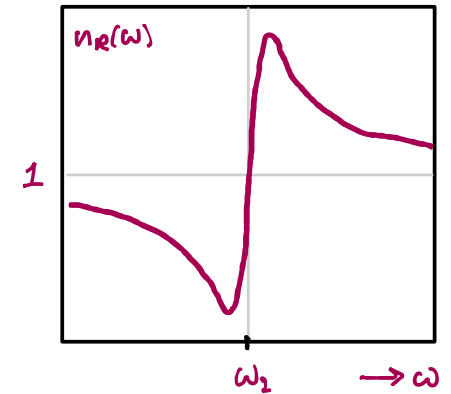
$$n_I = -\frac{a \omega}{2k} = -\frac{N \omega}{2k} \sigma(\Delta)$$

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$\omega < 1$  absorption



$\omega > 1$  gain



# Maxwell-Bloch Equations - Solitons

## Self-Induced Transparency & Solitons

- (\*) Example of a non-trivial application of the MBE's in the context of pulse propagation (highly dynamic, non-steady state behavior).
- (\*) The pulse area theorem suggests a light pulse with the proper envelope will act as a  $2\pi$  pulse. Thus, if the pulse is shorter than the excited state lifetime it may propagate without loss. Correct shaping may also allow propagation without changes in pulse shape.
- (\*) See Lecture Notes, Slusher & Gibbs 1972.

Envelope:  $\mathcal{E}(z,t) = \frac{2\hbar}{\mu\tau} \text{sech}(\xi/\tau)$ ,  $\xi = t - z/v$ ,  $\Delta = 0$

$$\Rightarrow \chi(z,t) = \frac{2}{v} \text{sech}(\xi/\tau), \quad \theta = \int_{-\infty}^{\infty} \chi(\xi/\tau) dt = 2\pi$$

Self-consistent solution with the the properties of a Soliton

$$\mathcal{E}(z,t) = \frac{2\hbar}{\mu\tau} \text{sech}(\xi/\tau)$$

$$\mu(\xi/\tau) = 0$$

$$v(\xi/\tau) = 2 \text{sech}(\xi/\tau) \tanh(\xi/\tau)$$

$$\omega(\xi/\tau) = -1 + 2 \text{sech}(\xi/\tau)$$

In the SVEA version of the Wave Eq.

$$\left(\frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t}\right) \mathcal{E}(z,t) = \frac{i\hbar}{\epsilon_0} N \mu^* (\mu - i\nu)$$

Substitute solutions for  $\mathcal{E}$ ,  $\mu$  and  $\nu$  to get

$$\frac{2\hbar}{\mu\tau} \left(\frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t}\right) \text{sech}\left(\frac{t-zv}{\tau}\right) =$$

$$\frac{2\hbar}{\mu\tau} \left(\frac{-1}{v\tau} + \frac{1}{c\tau}\right) \left[-\text{sech}\left(\frac{t-zv}{\tau}\right) \tanh\left(\frac{t-zv}{\tau}\right)\right] =$$

$$\frac{2\hbar N \mu^*}{\epsilon_0} \text{sech}\left(\frac{t-zv}{\tau}\right) \tanh\left(\frac{t-zv}{\tau}\right)$$

Solve for  $c/v$  to get

$$\frac{c}{v} = 1 + \frac{\hbar N |\mu|^2}{2\epsilon_0 \hbar} c\tau^2 = 1 + \frac{1}{2} a\beta c\tau^2$$

where  $a = \frac{\hbar N |\mu|^2}{2\epsilon_0 \hbar \beta} = N\sigma(\omega)$  (on-resonance absorption coeff.)

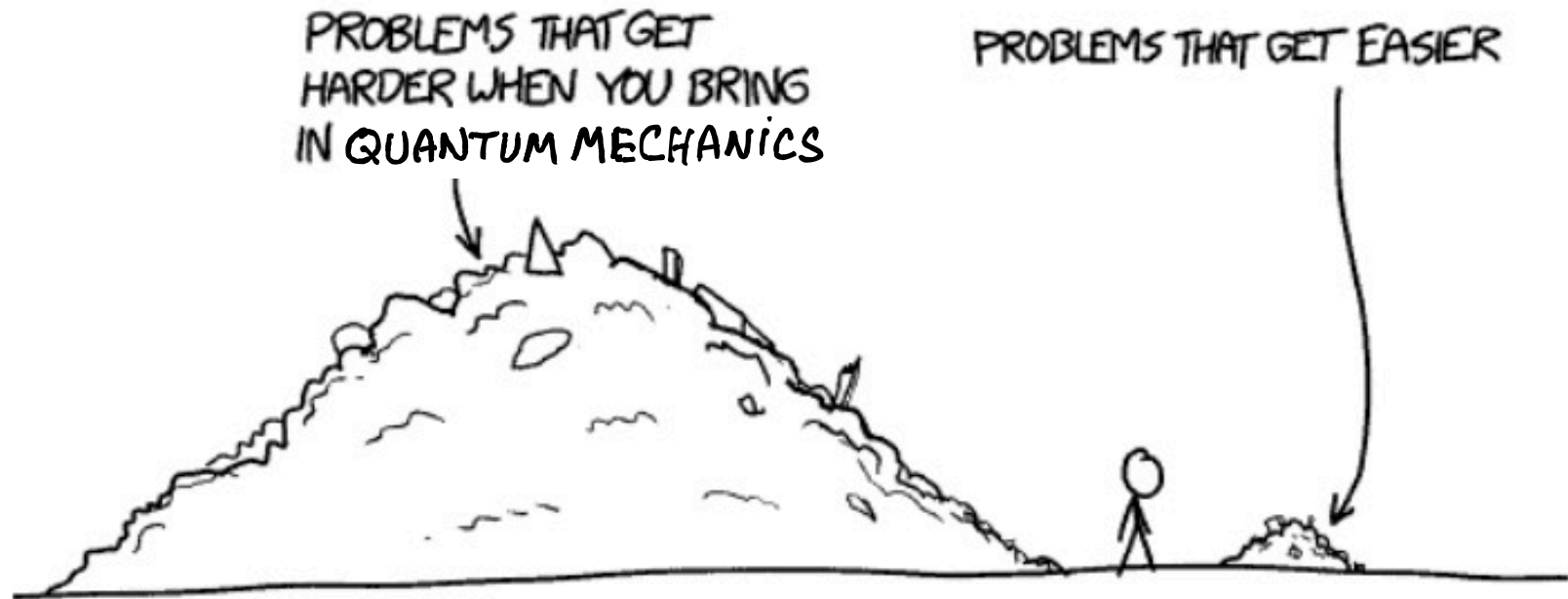
Consider Na vapor,  $\lambda = 589 \text{ nm}$ ,  $N = 10^{19} \text{ m}^{-3}$ ,  $T \sim 0 \text{ K}$ , and  $\beta = 2\pi \times 4.9 \text{ MHz}$  (completely opaque on res.)

Assuming  $v \sim \frac{1}{2}c \Rightarrow \frac{c}{v} \sim 2 = 1 + a\beta c\tau^2 \Rightarrow \frac{1}{2} a\beta c\tau^2 \sim 1$

we must have  $\tau \sim \sqrt{\frac{2}{a\beta c}} \sim 36 \text{ ps} \ll 16 \text{ ns} !!$



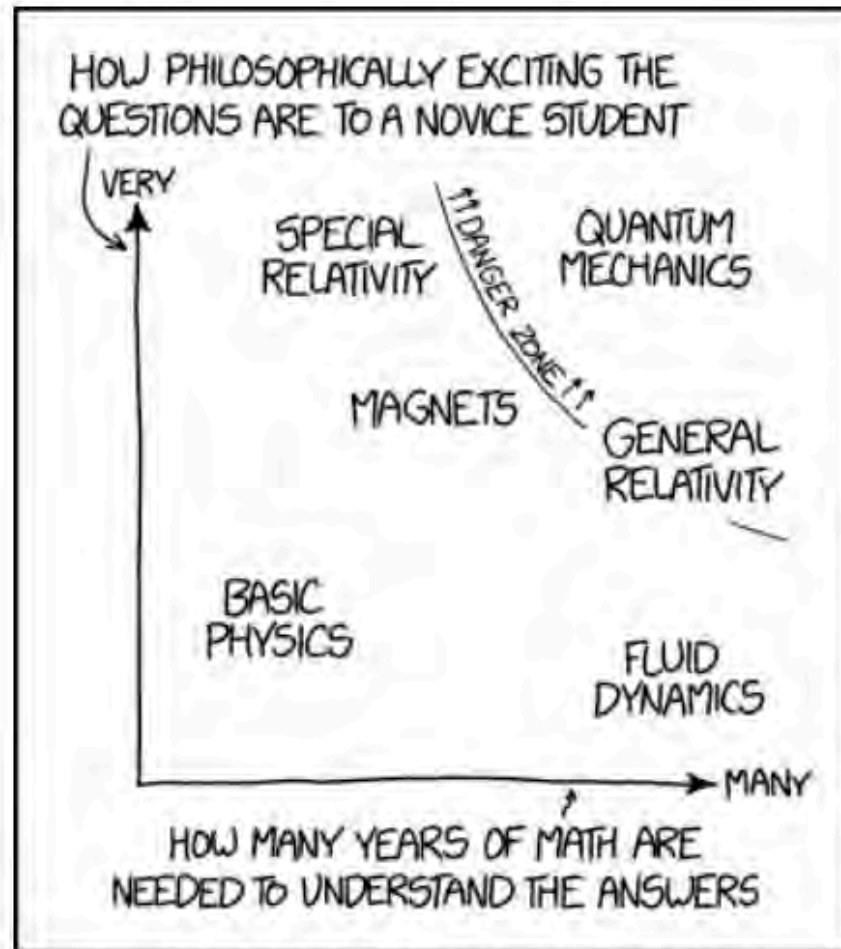
# Quantum Electrodynamics – Intro to Field Theory



Source: xkcd.com

# Quantum Electrodynamics – Intro to Field Theory

## QUANTUM



WHY SO MANY PEOPLE HAVE WEIRD IDEAS ABOUT QUANTUM MECHANICS

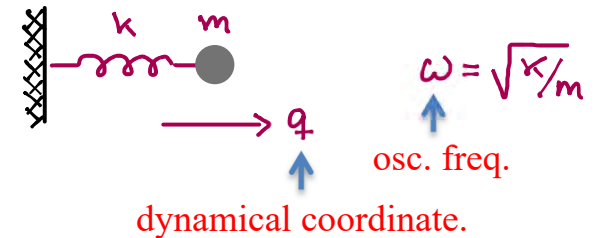
# Quantum Electrodynamics – Intro to Field Theory

- (\*) Primary goal of OPTI 544:  
Quantum description of EM field
- (\*) Challenge: 1st semester Grad level QM (OPTI 570) does not tell how to do this
- (\*) Warm-up: Quantum field theory for vibrations (sound) in elastic rod
- (\*) This is in part a review of the classical Lagrange/Hamilton-Jacobi description of continuous systems
- (\*) Here we present the formalism as a Cookbook Recipe for how we get from Classical to Quantum Physics

See, e. g., Cohen-Tannoudji Vol. 2,  
Appendix III, Sections 1-3.

## Classical Simple Harmonic Oscillator (SHO)

Particle on  
a spring



Kinetic Energy:

$$T = \frac{1}{2} m \dot{q}^2$$

Potential Energy:

$$V = \frac{1}{2} k q^2 = \frac{1}{2} m \omega^2 q^2$$

Lagrangian:

$$\mathcal{L} = T - V = \frac{1}{2} m \dot{q}^2 - \frac{1}{2} m \omega^2 q^2$$

$$\frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{q}} - \frac{\partial \mathcal{L}}{\partial q} = 0 \quad \Rightarrow \quad \ddot{q} + \omega^2 q = 0$$

usual eq. of motion

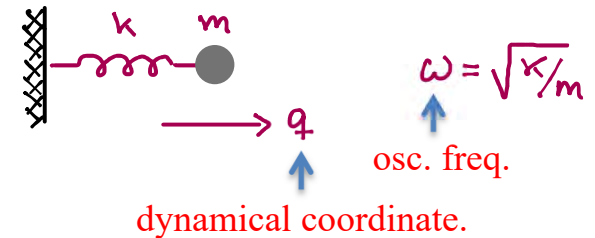
# Quantum Electrodynamics – Intro to Field Theory

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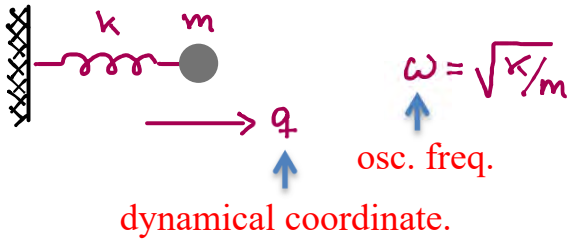
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# Quantum Electrodynamics – Intro to Field Theory

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usual eq. of motion

Conjugate momentum

$$p = \frac{\partial \mathcal{L}}{\partial \dot{q}} = m \dot{q}$$

Hamiltonian

$$\mathcal{H} = T(\dot{q} = p/m) + V(q) = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 q^2$$

$$\dot{q} = \frac{\partial \mathcal{H}}{\partial p} = p/m$$

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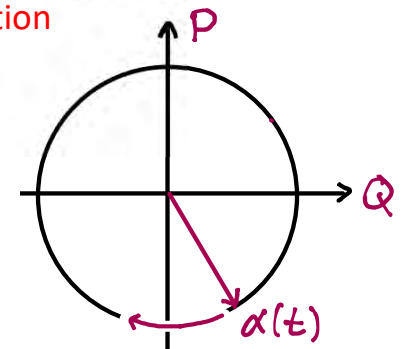
$$\ddot{q} + \omega^2 q = 0$$

Phase plane

Scaled variables

$$Q \equiv q/q_0, \quad P = p/p_0$$

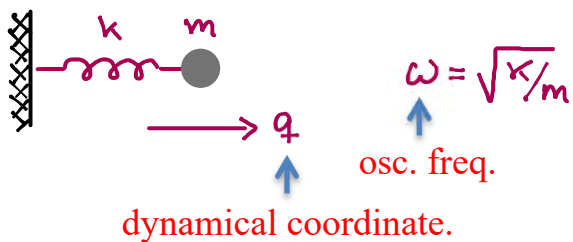
$$\alpha = Q + iP \begin{cases} Q = \text{Re}[\alpha] \\ P = \text{Im}[\alpha] \\ \mathcal{H} = E_0 \alpha^* \alpha \end{cases}$$



# Quantum Electrodynamics – Intro to Field Theory

## Classical Simple Harmonic Oscillator (SHO)

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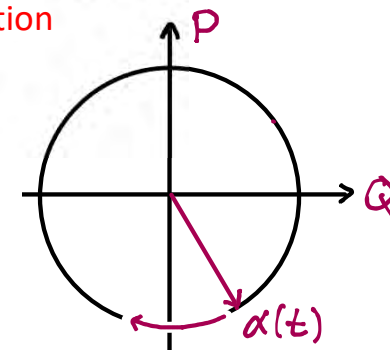
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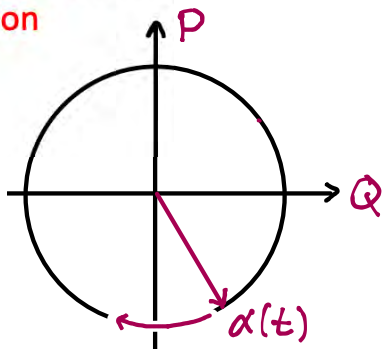
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## Quantum Harmonic Oscillator

Formal Quantization Procedure:

$$q \rightarrow \hat{q}, \quad p \rightarrow \hat{p}, \quad [\hat{q}, \hat{p}] = i\hbar$$

Choose  $E_0 = \hbar\omega \rightarrow q_0 = \sqrt{\frac{2\hbar}{m\omega}}, \quad p_0 = \sqrt{2m\hbar\omega}$

natural scale

$$\alpha \rightarrow \hat{a} = \hat{Q} + i\hat{P} = \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{q} + i\frac{\hat{p}}{m\omega} \right)$$

$$[\hat{a}, \hat{a}^\dagger] = 1$$

Rewrite:

$$\hat{H} = \hbar\omega (\hat{Q}^2 + \hat{P}^2) = \hbar\omega (\hat{a}^\dagger \hat{a} + 1/2)$$

$$\hat{N} = \hat{a}^\dagger \hat{a} \quad (\text{number operator})$$

# Quantum Electrodynamics – Intro to Field Theory

## Quantum Harmonic Oscillator

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$$\text{Commutator } [\hat{H}, \hat{N}] = 0$$

$\rightarrow$  joint energy/number states  $|n\rangle$

$$\hat{H}|n\rangle = \hbar\omega (n + \frac{1}{2})|n\rangle$$

$$\hat{N}|n\rangle = n|n\rangle$$

Commutators

$$\left. \begin{aligned} [\hat{N}, \hat{a}^\dagger] &= \hat{a}^\dagger \\ [\hat{N}, \hat{a}] &= -\hat{a} \end{aligned} \right\}$$



$$\hat{a}|n\rangle = \sqrt{n}|n-1\rangle$$

$$\hat{a}^\dagger|n\rangle = \sqrt{n+1}|n+1\rangle$$

$$\hat{a}|0\rangle = 0$$

Generating excited states

$$|n\rangle = \frac{1}{\sqrt{n!}} (\hat{a}^\dagger)^n |0\rangle$$

# Quantum Electrodynamics – Intro to Field Theory

## Quantum Harmonic Oscillator

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# Quantum Electrodynamics – Intro to Field Theory

Commutator  $[\hat{H}, \hat{N}] = 0$

→ joint energy/number states  $|n\rangle$

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Commutators

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Generating excited states

$$|n\rangle = \frac{1}{\sqrt{n!}} (\hat{a}^\dagger)^n |0\rangle$$

Expectation values for  $\hat{q}$  and  $\hat{p}$  in number states

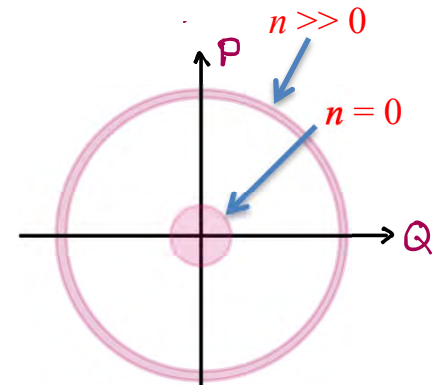
$$\langle n|\hat{q}|n\rangle = \langle n|\hat{p}|n\rangle = 0$$

$$\langle n|\hat{q}^2|n\rangle = \frac{q_0^2}{2}(n+1/2) \neq 0$$

$$\langle n|\hat{p}^2|n\rangle = \frac{p_0^2}{2}(n+1/2) \neq 0$$

$$\Delta q \Delta p = \frac{q_0 p_0}{2}(n+1/2) = \hbar(n+1/2)$$

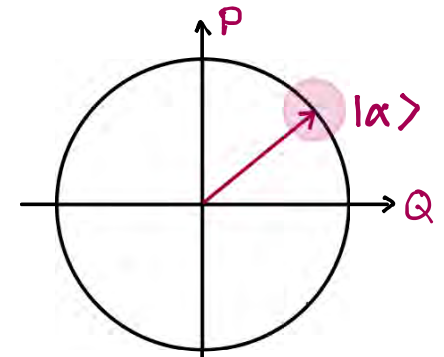
Phase space visualization of number states



Quasi-classical (coherent) state

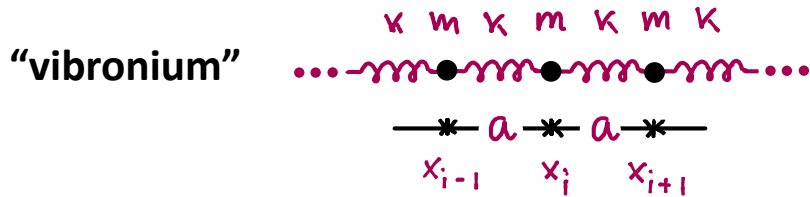
$$|\alpha\rangle = e^{-|\alpha|^2/2} \sum_i \frac{\alpha^i}{\sqrt{i!}} |i\rangle$$

$$\Delta q \Delta p = \hbar/2, \quad \Delta Q = \Delta P$$



# Quantum Electrodynamics – Intro to Field Theory

## Lagrange formulation of 1D Scalar Field



Configuration space =  $\{x_i\}$  (set of  $N$  osc. positions)

$$T = \sum_{i=1}^N \frac{1}{2} m \dot{x}_i^2, \quad V = \sum_{i=1}^N \frac{1}{2} \kappa (x_{i+1} - x_i)^2$$

Lagrangian, equations of motion

Continuum limit  $\rightarrow$  Elastic rod

$$\begin{aligned}
 N \rightarrow \infty & \quad m/a \rightarrow \mu \quad \leftarrow \text{linear mass density} \\
 a \rightarrow dx & \quad \kappa a \rightarrow \gamma \quad \leftarrow \text{Youngs modulus} \\
 \{x_i\} & \rightarrow \eta(x) \quad \leftarrow \text{displacement field (sound)}
 \end{aligned}$$

Rewrite

$$T = \lim_{N \rightarrow \infty} \sum_{i=1}^N a \frac{1}{2} \left(\frac{m}{a}\right) \dot{x}_i^2 = \int dx \frac{1}{2} \mu \left(\frac{\partial \eta}{\partial t}\right)^2$$

$$V = \lim_{N \rightarrow \infty} \sum_{i=1}^N a \frac{1}{2} \kappa a \left(\frac{x_{i+1} - x_i}{a}\right)^2 = \int dx \frac{1}{2} \gamma \left(\frac{\partial \eta}{\partial x}\right)^2$$

Lagrangian:

$$\mathcal{L} = T - V = \int dx \frac{1}{2} \mu \left(\frac{\partial \eta}{\partial t}\right)^2 - \int dx \frac{1}{2} \gamma \left(\frac{\partial \eta}{\partial x}\right)^2$$

Notes, Homework  $\rightarrow$  Scalar wave equation

$$\frac{\partial^2 \eta}{\partial t^2} - \frac{\gamma}{\mu} \frac{\partial^2 \eta}{\partial x^2} = 0$$

– Not yet ready for Quantization –

# Quantum Electrodynamics – Intro to Field Theory

Rewrite

$$T = \lim_{N \rightarrow \infty} \sum_{i=1}^N a \frac{1}{2} \left( \frac{m}{a} \dot{x}_i \right)^2 = \int dx \frac{1}{2} \mu \left( \frac{\partial \eta}{\partial t} \right)^2$$

$$V = \lim_{N \rightarrow \infty} \sum_{i=1}^N a \frac{1}{2} k a \left( \frac{x_{i+1} - x_i}{a} \right)^2 = \int dx \frac{1}{2} \gamma \left( \frac{\partial \eta}{\partial x} \right)^2$$

Lagrangian:

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## Normal Mode Decomposition

Field in cavity:



Solutions to wave eq.

Let  $\eta(x,t) = g(t)u(x) = g_0 e^{i\omega t} u(x) \rightarrow$

$$\ddot{g} - v^2 g'' = -\omega^2 g(t) u(x) - v^2 g(t) u''(x) = 0$$



$$u''(x) = -k^2 u(x), \quad k = \omega/v$$

Solutions in cavity:

$$u_k(x) = \sqrt{\frac{2}{L}} \sin(kx), \quad k = \frac{n\pi}{L}$$

These standing waves are a set of Normal Modes



# Quantum Electrodynamics – Intro to Field Theory

## Normal Mode Decomposition

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These modes are orthonormal and complete



$$\eta(x,t) = \sqrt{L} \sum_k g_k(t) u_k(x)$$

Normal mode expansion of  $\eta(x,t)$  in basis  $u_k(x)$

Lagrangian for the acoustic field:

$$\begin{aligned} T &= \int dx \frac{1}{2} \rho \left( \frac{\partial \eta}{\partial t} \right)^2 = \sum_{k,k'} \underbrace{\frac{1}{2} \rho L}_{M} \dot{q}_k \dot{q}_{k'} \underbrace{\int dx u_k(x) u_{k'}(x)}_{\delta_{kk'}} \\ &= \sum_k \frac{1}{2} M \dot{q}_k^2 \\ V &= \int dx \frac{1}{2} \gamma \left( \frac{\partial \eta}{\partial x} \right)^2 = \sum_{k,k'} \frac{1}{2} \gamma L q_k q_{k'} \int dx \left( \frac{\partial u_k}{\partial x} \right) \left( \frac{\partial u_{k'}}{\partial x} \right) \\ &= \sum_k \frac{1}{2} M \omega_k^2 q_k^2 \end{aligned}$$

# Quantum Electrodynamics – Intro to Field Theory

## Normal Mode Decomposition

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# Quantum Electrodynamics – Intro to Field Theory

The rest now follows from the Lagrangian

$$\mathcal{L} = T - V = \sum_{\mathbf{k}} \left( \frac{1}{2} M \dot{q}_{\mathbf{k}}^2 - \frac{1}{2} M \omega_{\mathbf{k}}^2 q_{\mathbf{k}}^2 \right) = \sum_{\mathbf{k}} \mathcal{L}_{\mathbf{k}}$$



Canonical  
Momentum

$$p_{\mathbf{k}} = \frac{\partial \mathcal{L}}{\partial \dot{q}_{\mathbf{k}}} = M \dot{q}_{\mathbf{k}}$$

Hamiltonian

$$\mathcal{H}(\{p_{\mathbf{k}}, q_{\mathbf{k}}\}) = T + V = \sum_{\mathbf{k}} \left( \frac{p_{\mathbf{k}}^2}{2M} + \frac{1}{2} M \omega_{\mathbf{k}}^2 q_{\mathbf{k}}^2 \right)$$

(collection of SHO's, one for each normal mode)

Following the standard recipe...

$$E_{0,\mathbf{k}} = \hbar \omega_{\mathbf{k}}, \quad q_{0,\mathbf{k}} = \sqrt{2\hbar/M\omega_{\mathbf{k}}}, \quad p_{0,\mathbf{k}} = \sqrt{2M\hbar\omega_{\mathbf{k}}}$$

$$Q_{\mathbf{k}} = q_{\mathbf{k}}/q_{0,\mathbf{k}}, \quad P_{\mathbf{k}} = p_{\mathbf{k}}/p_{0,\mathbf{k}}, \quad \alpha_{\mathbf{k}} = Q_{\mathbf{k}} + iP_{\mathbf{k}}$$

... we get solutions

$$\alpha_{\mathbf{k}}(t) = Q_{\mathbf{k}}(t) + iP_{\mathbf{k}}(t) = \alpha_{\mathbf{k}}(0) e^{-i\omega_{\mathbf{k}}t}$$

This finally gives us

$$\mathcal{H} = \sum_{\mathbf{k}} \hbar \omega_{\mathbf{k}} (Q_{\mathbf{k}}^2 + P_{\mathbf{k}}^2) = \sum_{\mathbf{k}} \hbar \omega_{\mathbf{k}} \alpha_{\mathbf{k}}^* \alpha_{\mathbf{k}}$$

$$y(x, L) = \sqrt{L} \sum_{\mathbf{k}} q_{\mathbf{k}}(t) u_{\mathbf{k}}(x)$$

$$= \frac{1}{2} \sum_{\mathbf{k}} \sqrt{L q_{0,\mathbf{k}}^2} \left( \alpha_{\mathbf{k}}(t) u_{\mathbf{k}}(x) + \alpha_{\mathbf{k}}^*(t) u_{\mathbf{k}}^*(x) \right)$$