

# Light-Matter Interaction

## Hierarchy of Sophistication:

- Classical      Classical light, classical matter
- Semiclassical      Classical light, quantum matter
- Quantum      Quantum light, quantum matter

## Possible attitudes:

- Purist      Most complete description possible
- Minimalist      Quantum only when necessary
- Pragmatic      Quantum or classical, based on what is simplest and still works


**OPTI 544: All of the above in turn**

## Classical Theory of Light-Matter Interaction

Self-consistent, fully classical description



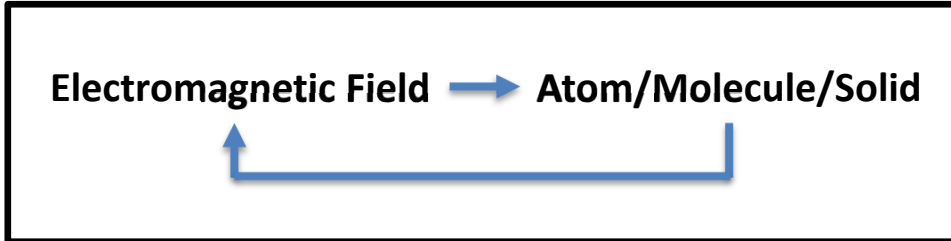
## Motivation: We will

- Develop Concepts  $\alpha(\omega), n, \chi$
- Develop Intuition
- Classical is often adequate, sometimes accurate
- A Quantum Theory has classical limits  Identify/understand regime of validity
- The Classical description is a useful starting point for Nonlinear and Quantum Optics

# Light-Matter Interaction

## Classical Theory of Light-Matter Interaction

Self-consistent, fully classical description



Motivation: We will

- Develop Concepts  $\alpha(\omega), n, \chi$
- Develop Intuition
- Classical is often adequate, sometimes accurate
- A Quantum Theory has classical limits  $\rightarrow$   
Identify/understand regime of validity
- The Classical description is a useful starting point for Nonlinear Optics

## The Electromagnetic Field: Basic Eqs. in SI Units

### Maxwell's eqs.

( no free charges, currents  $\rightarrow$  dielectrics )

- (i)  $\nabla \cdot \vec{D} = \rho = 0$        $\vec{D}$ : Dielectric displacement
- (ii)  $\nabla \cdot \vec{B} = 0$                $\vec{B}$ : Magnetic induction
- (iii)  $\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$        $\vec{E}$ : Electric field
- (iv)  $\nabla \times \vec{H} = \frac{\partial \vec{D}}{\partial t} + \vec{J}$        $\vec{H}$ : Magnetic field

### Material Response:

- (v)  $\vec{B} = \mu_0 \vec{H} + \vec{M}$        $\leftarrow$  Non-magnetic  $\rightarrow \vec{M} = 0$
- (vi)  $\vec{D} = \epsilon_0 \vec{E} + \vec{P}$        $\leftarrow$  Info about response in dipole moment density (polarization density)

# Light-Matter Interaction

## The Electromagnetic Field: Basic Eqs. in SI Units

### Maxwell's eqs.

( no free charges, currents → dielectrics )

(i)  $\nabla \cdot \vec{D} = \rho = 0$        $\vec{D}$ : Dielectric displacement

(ii)  $\nabla \cdot \vec{B} = 0$        $\vec{B}$ : Magnetic induction

(iii)  $\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$        $\vec{E}$ : Electric field

(iv)  $\nabla \times \vec{H} = \frac{\partial \vec{D}}{\partial t} + \vec{J}$        $\vec{H}$ : Magnetic field

### Material Response:

(v)  $\vec{B} = \mu_0 \vec{H} + \vec{M}$       ← Non-magnetic →  $\vec{M} = 0$

(vi)  $\vec{D} = \epsilon_0 \vec{E} + \vec{P}$       ← Info about response in dipole moment density (polarization density)

We need equations that describe:

- the behavior of  $\vec{E}$  for given  $\vec{P}$
- the medium response  $\vec{P}$  for given  $\vec{E}$

### Wave Equation:

Take curl of (iii), then use (iv)

$$\nabla \times (\nabla \times \vec{E}) = -\nabla \times \frac{\partial \vec{B}}{\partial t} = -\frac{\partial}{\partial t} (\nabla \times \vec{B}) = -\mu_0 \frac{\partial^2 \vec{D}}{\partial t^2}$$

Next, use the identity

$$\nabla \times (\nabla \times \vec{E}) = \nabla (\nabla \cdot \vec{E}) - \nabla^2 \vec{E}$$

to obtain  $\vec{D} = \nabla (\nabla \cdot \vec{E}) = -\mu_0 \frac{\partial^2 \vec{D}}{\partial t^2}$

Finally, let  $\vec{D} = \epsilon_0 \vec{E} + \vec{P}$  and use  $\epsilon_0 \mu_0 = \frac{1}{c^2}$

to obtain

$$-\nabla (\nabla \cdot \vec{E}) + \nabla^2 \vec{E} = \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} + \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \vec{P}}{\partial t^2}$$

This is the Wave Equation, still exact in this form

# Light-Matter Interaction

We need equations that describe:

- the behavior of  $\vec{E}$  for given  $\vec{p}$
- the medium response  $\vec{p}$  for given  $\vec{E}$

## Wave Equation:

Take curl of (iii), then use (iv)

$$\nabla \times (\nabla \times \vec{E}) = -\nabla \times \frac{\partial \vec{B}}{\partial t} = -\frac{\partial}{\partial t} (\nabla \times \vec{B}) = -\mu_0 \frac{\partial^2 \vec{D}}{\partial t^2}$$

Next, use the identity

$$\nabla \times (\nabla \times \vec{E}) = \nabla (\nabla \cdot \vec{E}) - \nabla^2 \vec{E}$$

to obtain 
$$\vec{D} = \nabla (\nabla \cdot \vec{E}) - \mu_0 \frac{\partial^2 \vec{D}}{\partial t^2}$$

Finally, let  $\vec{D} = \epsilon_0 \vec{E} + \vec{p}$  and use  $\epsilon_0 \mu_0 = \frac{1}{c^2}$

to obtain

$$-\nabla (\nabla \cdot \vec{E}) + \nabla^2 \vec{E} = \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} + \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \vec{p}}{\partial t^2}$$

This is the Wave Equation, still exact in this form

## Transverse Fields

Definition: a field for which  $\nabla \cdot \vec{E} = 0$  is Transverse

Example: a plane wave,  $\vec{E}(\vec{r}, t) = \vec{E}(t) e^{i\vec{k} \cdot \vec{r}}$ , where  $\vec{E}(t) \perp \vec{k}$ , is transverse.

The physical field is  $\text{Re}[\vec{E}(\vec{r}, t)]$

For transverse fields the wave equation simplifies to

$$\nabla^2 \vec{E} - \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \vec{p}}{\partial t^2}$$

This version of the wave equation can be a poor approximation in non-isotropic media!

# Light-Matter Interaction

## Transverse Fields

Definition: a field for which  $\nabla \cdot \vec{E} = 0$   
is Transverse

Example: a plane wave,  $\vec{E}(\vec{r}, t) = \vec{E}(t) e^{i\vec{k} \cdot \vec{r}}$ ,  
where  $\vec{E}(t) \perp \vec{k}$ , is transverse.

The physical field is  $\text{Re}[\vec{E}(\vec{r}, t)]$

For transverse fields the wave equation simplifies to

$$\nabla^2 \vec{E} - \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \vec{P}}{\partial t^2}$$

This version of the wave equation can be a poor approximation in non-isotropic media!

## Isotropic Media

Absent a preferred direction, the induced  $\vec{P}$   
must be parallel to the driving field  $\vec{E}$

Linear response, most general case:

$$\begin{aligned} \vec{D}(t) &= \epsilon_0 \vec{E}(t) + \vec{P}(t) \\ &= \epsilon_0 \vec{E}(t) + \epsilon_0 \int_{-\infty}^t dt' R(t-t') \vec{E}(t') \end{aligned}$$

where the *response function*  $R(t-t')$  is a scalar  
and we have  $R(\tau) = 0$  for  $\tau < 0$

Take divergence on both sides and use M.E. (i)

$$\begin{aligned} \nabla \cdot \vec{D}(t) &= \epsilon_0 \nabla \cdot \vec{E}(t) + \epsilon_0 \int_{-\infty}^t dt' R(t-t') \nabla \cdot \vec{E}(t') = 0 \\ \Rightarrow \nabla \cdot \vec{E}(t) &= - \int_{-\infty}^t dt' R(t-t') \nabla \cdot \vec{E}(t') \text{ for all } t \end{aligned}$$

It follows that  $\nabla \cdot \vec{E}(t) = 0$  for all  $t$ ,

OR  $R(\tau) = -2\delta(\tau)$

# Light-Matter Interaction

## Isotropic Media

Absent a preferred direction, the induced  $\vec{P}$  must be parallel to the driving field  $\vec{E}$

Linear response, most general case:

$$\begin{aligned}\vec{D}(t) &= \epsilon_0 \vec{E}(t) + \vec{P}(t) \\ &= \epsilon_0 \vec{E}(t) + \epsilon_0 \int_{-\infty}^t dt' R(t-t') \vec{E}(t')\end{aligned}$$

where the *response function*  $R(t-t')$  is a scalar and we have  $R(\tau) = 0$  for  $\tau < 0$

Take divergence on both sides and use M.E. (i)

$$\nabla \cdot \vec{D}(t) = \epsilon_0 \nabla \cdot \vec{E}(t) + \epsilon_0 \int_{-\infty}^t dt' R(t-t') \nabla \cdot \vec{E}(t') = 0$$

and  $\nabla \cdot \vec{E}(t) = - \int_{-\infty}^t dt' R(t-t') \nabla \cdot \vec{E}(t')$  for all  $t$

It follows that  $\nabla \cdot \vec{E}(t) = 0$  for all  $t$ ,

OR  $R(\tau) = -2\delta(\tau)$

Note: if  $R(\tau) \propto \delta(\tau)$  (instantaneous response) then

$$\epsilon_0 \int_{-\infty}^t dt' R(t-t') \vec{E}(t') = \epsilon_0 \chi \vec{E}(t)$$

↑ susceptibility

The case  $R(\tau) = -2\delta(\tau)$  is an example of negative susceptibility,  $\chi < 0$ , which only occurs in certain engineered metamaterials.



Electric fields are transverse in linear, isotropic dielectric media  
(including the vacuum)

Wave Equation in free space

$$\nabla^2 \vec{E} - \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} = 0$$

# Light-Matter Interaction

Note: if  $R(\tau) \propto \delta(\tau)$  (instantaneous response) then

$$\epsilon_0 \int_{-\infty}^t dt' R(t-t') \vec{E}(t') = \epsilon_0 \chi \vec{E}(t)$$

↑ susceptibility

The case  $R(\tau) = -2\delta(\tau)$  is an example of negative susceptibility,  $\chi < 0$ , which only occurs in certain engineered metamaterials.



Electric fields are transverse in linear, isotropic dielectric media  
(including the vacuum)

Wave Equation in free space

$$\nabla^2 \vec{E} - \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} = 0$$

Monochromatic trial solution  $\vec{E}(\vec{r}, t) = \vec{E}_0(\vec{r}) e^{-i\omega t}$



$$\nabla^2 \vec{E}_0(\vec{r}) e^{-i\omega t} + \frac{\omega^2}{c^2} \vec{E}_0(\vec{r}) e^{-i\omega t} = 0$$

Equation for the spatial component alone:

$$\nabla^2 \vec{E}_0(\vec{r}) + |\vec{k}|^2 \vec{E}_0(\vec{r}) = 0, \quad |\vec{k}| = \omega/c$$



Plane wave solutions

$$\vec{E}_0(\vec{r}) = \vec{E} E_0 e^{i\vec{k} \cdot \vec{r}}, \quad |\vec{k}| = \omega/c$$

**Optical Cavities:** Here we need to solve the wave equation subject to boundary conditions. See, e. g., Millony & Eberly for examples such as rectangular cavities, Fabryt-Perot etalons, and spherical mirror resonators.

# Light-Matter Interaction

Monochromatic trial solution  $\vec{E}(\vec{r}, t) = \vec{E}_0(\vec{r}) e^{-i\omega t}$



$$\nabla^2 \vec{E}_0(\vec{r}) e^{-i\omega t} + \frac{\omega^2}{c^2} \vec{E}_0(\vec{r}) e^{-i\omega t} = 0$$

Equation for the spatial component alone:

$$\nabla^2 \vec{E}_0(\vec{r}) + |\vec{k}|^2 \vec{E}_0(\vec{r}) = 0, \quad |\vec{k}| = \omega/c$$



Plane wave solutions

$$\vec{E}_0(\vec{r}) = \vec{\epsilon} E_0 e^{i\vec{k} \cdot \vec{r}}, \quad |\vec{k}| = \omega/c$$

**Optical Cavities:** Here we need to solve the wave equation subject to boundary conditions. See, e. g., Millony & Eberly for examples such as rectangular cavities, Fabryt-Perot etalons, and spherical mirror resonators.

Wave Equation in Fourier Space:

In Configuration Space:

$$\nabla^2 \vec{E}(\vec{r}, t) - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \vec{E}(\vec{r}, t) = \frac{1}{\epsilon_0 c^2} \frac{\partial^2}{\partial t^2} \vec{P}(\vec{r}, t)$$



In Fourier Space:

$$k^2 \vec{E}(\vec{k}, \omega) - \frac{\omega^2}{c^2} \vec{E}(\vec{k}, \omega) = \frac{\omega^2}{\epsilon_0 c^2} \vec{P}(\vec{k}, \omega)$$

**Note:** In the Fourier domain the wave equation is purely algebraic – there are no derivatives or integrals. This becomes important later in the course when we quantize the electromagnetic field.



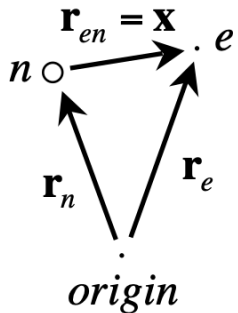
# Light-Matter Interaction

## Theory of Atomic Response

So far, we have a model for the field. Next, we need a model of how the constituents of the medium responds to the field.

This will allow us to find the polarization density  $\vec{P}$  as function of the field  $\vec{E}$

Classical "atom"



Simple model:  
nucleus + electron

Lorentz Force

$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

$\sim 0$  if non-relativistic

Newton:

(i)  $m_n \frac{d^2}{dt^2} \vec{r}_n(t) = -e \vec{E}(\vec{r}_n, t) - \vec{F}_{en}(\vec{r}_{en}, t)$

(ii)  $m_e \frac{d^2}{dt^2} \vec{r}_e(t) = e \vec{E}(\vec{r}_e, t) + \vec{F}_{en}(\vec{r}_{en}, t)$

This is a standard 2-body problem which we can re-cast as in terms of relative and COM motion.

We define:

$$\vec{x} = \vec{r}_{en} = \vec{r}_e - \vec{r}_n \qquad m = \frac{m_e m_n}{m_e + m_n} \sim m_e$$

$$\vec{R} = \frac{m_e \vec{r}_e + m_n \vec{r}_n}{M} \qquad M = m_e + m_n \sim m_n$$

$\vec{x}$  Relative coord.       $m$  Reduced mass

$\vec{R}$  Center-of-mass       $M$  Total mass

# Light-Matter Interaction

Newton:

$$(i) \quad m_n \frac{d^2}{dt^2} \vec{r}_n(t) = -e \vec{E}(\vec{r}_n, t) - \vec{F}_{en}(\vec{r}_{en}, t)$$

$$(ii) \quad m_e \frac{d^2}{dt^2} \vec{r}_e(t) = e \vec{E}(\vec{r}_e, t) + \vec{F}_{en}(\vec{r}_{en}, t)$$

This is a standard 2-body problem which we can re-cast as in terms of relative and COM motion.

We define:

$$\vec{x} = \vec{r}_{en} = \vec{r}_e - \vec{r}_n \quad m = \frac{m_e m_n}{m_e + m_n} \sim m_e$$

$$\vec{R} = \frac{m_e \vec{r}_e + m_n \vec{r}_n}{M} \quad M = m_e + m_n \sim m_n$$

$\vec{x}$  Relative coord.       $m$  Reduced mass

$\vec{R}$  Center-of-mass       $M$  Total mass

Sub into (i), (ii) and rewrite:

$$M \frac{d^2}{dt^2} \vec{R} = e \left[ \vec{E} \left( \vec{R} + \frac{m_n}{M} \vec{x}, t \right) - \vec{E} \left( \vec{R} - \frac{m_e}{M} \vec{x}, t \right) \right]$$

$$m \frac{d^2}{dt^2} \vec{x} = \frac{e}{2} \left[ \vec{E} \left( \vec{R} + \frac{m_n}{M} \vec{x}, t \right) + \vec{E} \left( \vec{R} - \frac{m_e}{M} \vec{x}, t \right) \right] \\ + \vec{F}_{en}(\vec{x}) + \frac{1}{2} (m_n - m_e) \frac{d^2}{dt^2} \vec{R}$$

Basic result, no approximations !

Milloni & Eberly,  
main text



Set  $\vec{R} \approx \vec{r}_n$ ,  $\vec{x} \approx \vec{r}_{en}$   
Throw away eq. for  $\vec{R}$

Electric Dipole approximation

Atomic dimensions      Optical Wavelength

$$|\vec{x}| \sim 1 \text{ \AA} \quad \ll \quad \lambda \sim 10^4 \text{ \AA}$$

# Light-Matter Interaction

Sub into (i), (ii) and rewrite:

$$M \frac{d^2}{dt^2} \vec{R} = e \left[ \vec{E} \left( \vec{R} + \frac{m_n}{M} \vec{x}, t \right) - \vec{E} \left( \vec{R} - \frac{m_e}{M} \vec{x}, t \right) \right]$$

$$m \frac{d^2}{dt^2} \vec{x} = \frac{e}{2} \left[ \vec{E} \left( \vec{R} + \frac{m_n}{M} \vec{x}, t \right) + \vec{E} \left( \vec{R} - \frac{m_e}{M} \vec{x}, t \right) \right] + \vec{F}_{en}(\vec{x}) + \frac{1}{2} (m_n - m_e) \frac{d^2}{dt^2} \vec{R}$$

Basic result, no approximations !

Milloni & Eberly,  
main text



Set  $\vec{R} \approx \vec{r}_n$ ,  $\vec{x} \approx \vec{r}_{en}$   
Throw away eq. for  $\vec{R}$

## Electric Dipole approximation

Atomic dimensions      Optical Wavelength

$$|\vec{x}| \sim 1 \text{ \AA} \quad \ll \quad \lambda \sim 10^4 \text{ \AA}$$

**EDA:** the field is nearly constant on the scale of an atom

Good approximation: 1<sup>st</sup> order expansion in  $\vec{x}$



$$\vec{E} \left( \vec{R} - \frac{m_e}{M} \vec{x}, t \right) \approx \vec{E}(\vec{R}, t) - \frac{m_e}{M} (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t)$$

$$\vec{E} \left( \vec{R} + \frac{m_n}{M} \vec{x}, t \right) \approx \vec{E}(\vec{R}, t) + \frac{m_n}{M} (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t)$$



$$M \frac{d^2}{dt^2} \vec{R} = e (\vec{x} \cdot \nabla) E(\vec{R}, t) \quad \text{COM}$$

$$m \frac{d^2}{dt^2} \vec{x} = e \vec{E}(\vec{R}, t) + \frac{m_n - m_e}{M} e (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t) + \vec{F}_{en}(\vec{x}) \quad \text{Rel. Coord.}$$

# Light-Matter Interaction

**EDA:** the field is nearly constant on the scale of an atom

Good approximation: 1<sup>st</sup> order expansion in  $\vec{x}$



$$\vec{E}(\vec{R} - \frac{m_e}{M} \vec{x}, t) \approx \vec{E}(\vec{R}, t) - \frac{m_e}{M} (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t)$$

$$\vec{E}(\vec{R} + \frac{m_e}{M} \vec{x}, t) \approx \vec{E}(\vec{R}, t) + \frac{m_e}{M} (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t)$$



$$M \frac{d^2}{dt^2} \vec{R} \approx e (\vec{x} \cdot \nabla) E(\vec{R}, t) \quad \text{COM}$$

$$m \frac{d^2}{dt^2} \vec{x} = e \vec{E}(\vec{R}, t) + \frac{m_n - m_e}{M} e (\vec{x} \cdot \nabla) \vec{E}(\vec{R}, t) + \vec{F}_{en}(\vec{x}) \quad \text{Rel. Coord.}$$

Physical Interpretation:

$\vec{p} = e\vec{x}$ : electric dipole moment of the atom



The Eqs. Of motion can then be recast as

$$M \frac{d^2}{dt^2} \vec{R} \approx (\vec{p} \cdot \nabla) \vec{E}(\vec{R}, t) = \vec{F} = -\nabla_{\vec{R}} V(\vec{x}, \vec{R}, t)$$

$$m \frac{d^2}{dt^2} \vec{x} = e \vec{E}(\vec{R}, t) + \vec{F}_{en}(\vec{x}) = -\nabla_{\vec{x}} V(\vec{x}, \vec{R}, t)$$

where  $V(\vec{x}, \vec{R}, t) = -\vec{p} \cdot \vec{E}(\vec{R}, t)$



electric-dipole interaction

**Note:** The COM Eq. is the foundation for a range of laser Atom Traps and Optical Tweezers. We will not explore this further in OPTI 544 lectures, but good review articles can be found in the published literature.

# Light-Matter Interaction

Physical Interpretation:

$\vec{p} = e\vec{x}$  : electric dipole moment of the atom



The Eqs. Of motion can then be recast as

$$M \frac{d^2}{dt^2} \vec{R} \approx (\vec{p} \cdot \nabla) \vec{E}(\vec{R}, t) = \vec{F} = -\nabla_{\vec{R}} V(\vec{x}, \vec{R}, t)$$

$$m \frac{d^2}{dt^2} \vec{x} = e \vec{E}(\vec{R}, t) + \vec{F}_{en}(\vec{x}) = -\nabla_{\vec{x}} V(\vec{x}, \vec{R}, t)$$

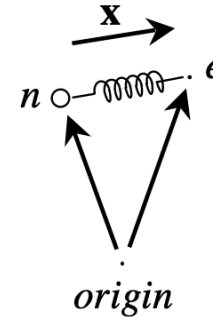
where  $V(\vec{x}, \vec{R}, t) = -\vec{p} \cdot \vec{E}(\vec{R}, t)$

electric-dipole interaction

Note: The COM Eq. is the foundation for a range of laser Atom Traps and Optical Tweezers. We will not explore this further in OPTI 544 lectures, but good review articles can be found in the published literature.

## The Electron Oscillator/Lorentz Oscillator

Simple model w/a harmonically bound electron:

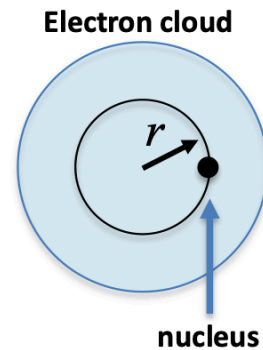


$$\vec{F}_{en} = -m\omega_0^2 \vec{x}$$

↑  
resonance frequency

This is meant as a model of the atomic response, not a model of the atom itself.

Nevertheless: QM suggest the atom consists of a point-like nucleus and a spherical electron cloud



Force from charge inside  $r$  as if entire charge was at the center

Force from charge outside  $r$  is zero

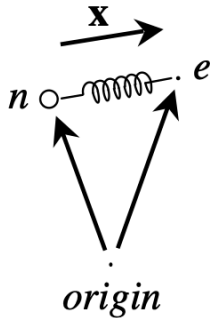
$$\text{Force } F \propto \frac{r^3}{r^2} \propto r$$

↑  
harmonic restoring force

# Light-Matter Interaction

## The Electron Oscillator/Lorentz Oscillator

Simple model w/a harmonically bound electron:

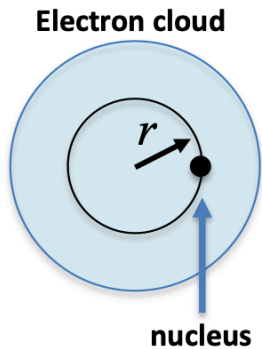


$$\vec{F}_{en} = -m\omega_0^2 \vec{x}$$

↑  
resonance frequency

This is meant as a model of the atomic response, not a model of the atom itself.

Nevertheless: QM suggest the atom consists of a point-like nucleus and a spherical electron cloud



Force from charge inside  $r$  as if entire charge was at the center

Force from charge outside  $r$  is zero

$$\text{Force } F \propto \frac{r^3}{r^2} \propto r$$

↑  
harmonic restoring force

Now substitute  $\vec{F}_{en} = -m\omega_0^2 \vec{x}$  into eq. for  $\vec{x}$



$$\frac{d^2}{dt^2} \vec{x} + \omega_0^2 \vec{x} = \frac{e}{m} \vec{E}(\vec{r}, t)$$

Combine with  $\vec{P} = N\vec{p}$ ,  $\vec{p} = e\vec{x}$  where  $N$  is the number density of atoms. This relates the macroscopic  $\vec{P}$  to the microscopic  $\vec{x}$

We now have

Maxwell's Equations  
The Lorentz model



Maxwell-Lorentz Equations  
We can seek self-consistent solutions to wave propagation

# Light-Matter Interaction

Now substitute  $\vec{F}_{en} = -m\omega_0^2 \vec{x}$  into eq. for  $\vec{x}$



$$\frac{d^2}{dt^2} \vec{x} + \omega_0^2 \vec{x} = \frac{e}{m} \vec{E}(\vec{r}, t)$$

Combine with  $\vec{P} = N\vec{p}$ ,  $\vec{p} = e\vec{x}$  where  $N$  is the number density of atoms. This relates the macroscopic  $\vec{P}$  to the microscopic  $\vec{x}$

We now have

Maxwell's Equations  
The Lorentz model



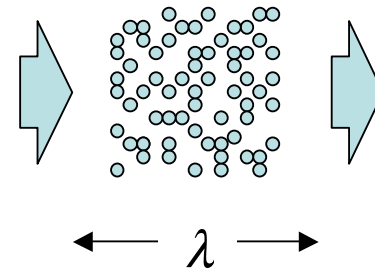
Maxwell-Lorentz Equations  
We can seek self-consistent solutions to wave propagation

## Classical Model of Absorption & Dispersion

Maxwell's Eqs: Oscillating dipole loses energy

➔ Must include damping in Eq, of Motion

Note: In perfectly homogeneous media the coherently scattered light from a collection of Lorentz oscillators interferes constructively only in the forward direction ➔



No energy loss for a propagating fields  
(See note set "Classical Light-Matter")

QM to the rescue: Part of the radiation from quantum mechanical atoms is incoherent.

For now we add damping "by hand"

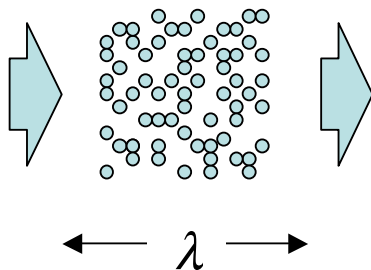
# Light-Matter Interaction

## Classical Model of Absorption & Dispersion

Maxwell's Eqs: Oscillating dipole loses energy

➔ Must include damping in Eq, of Motion

**Note:** In perfectly homogeneous media the coherently scattered light from a collection of Lorentz oscillators interferes constructively only in the forward direction ➔



No energy loss for a propagating fields  
(See note set "Classical Light-Matter")

QM to the rescue: Part of the radiation from quantum mechanical atoms is incoherent.

For now we add damping "by hand"

## The Lorentz Model with Damping

We add an ad hoc friction term  $w/\beta \ll \omega_0$   
↑  
 damping rate

This gives us our basic equation for the atomic response:

$$\frac{d^2}{dt^2} \vec{x} + 2\beta \frac{d}{dt} \vec{x} + \omega_0^2 \vec{x} = \frac{e}{m} \vec{E}(\vec{R}, t)$$

This type of differential equation generally has both oscillating and decaying terms. Solutions without source terms generally decay as  $e^{-\beta t}$

➔ We adopt a trial solution

Driving Field	$E(\vec{R}, t) = \vec{\epsilon} E_0 e^{-i(\omega t - k z)}$
Response	$\vec{x}(\vec{R}, t) = \vec{a} e^{-i(\omega t - k z)}$
	↑
	complex amplitude



# Light-Matter Interaction

## The Lorentz Model with Damping

We add an ad hoc friction term  $w/\beta \ll \omega_0$   
 $\uparrow$   
 damping rate

This gives us our basic equation for the atomic response:

$$\frac{d^2}{dt^2} \vec{x} + 2\beta \frac{d}{dt} \vec{x} + \omega_0^2 \vec{x} = \frac{e}{m} \vec{E}(\vec{R}, t)$$

This type of differential equation generally has both oscillating and decaying terms. Solutions without source terms generally decay as  $e^{-\beta t}$

➔ We adopt a trial solution

Driving Field  $\vec{E}(\vec{R}, t) = \vec{\Sigma} E_0 e^{-i(\omega t - k z)}$

Response  $\vec{x}(\vec{R}, t) = \vec{a} e^{-i(\omega t - k z)}$

$\uparrow$   
 complex amplitude

Solution for  $\vec{a}$  :

$$\vec{a} = -\vec{\Sigma} \frac{(e/m) E_0}{\omega^2 - \omega_0^2 + 2i\beta\omega}$$

Physical Quantities:

Field

$$\text{Re}[\vec{E}(\vec{R}, t)] = \vec{\Sigma} E_0 \cos(\omega t)$$

Dipole ( $\vec{\Sigma}$  real)

$$\text{Re}[\vec{p}(\vec{R}, t)] = \text{Re}[e\vec{x}(\vec{R}, t)]$$

$$= \vec{\Sigma} E_0 \frac{e^2}{m} \frac{(\omega_0^2 - \omega^2) \cos(\omega t - k z) + 2\beta\omega \sin(\omega t - k z)}{(\omega_0^2 - \omega^2) + 4\beta^2\omega^2}$$

Note:  $\vec{p}$  and  $\vec{E}$  generally oscillate out of phase

$\omega \ll \omega_0$  ➔  $\vec{p}$  &  $\vec{E}$  in-phase

$\omega = \omega_0$  ➔  $\vec{p}$  lags  $\vec{E}$  by  $\pi/2$

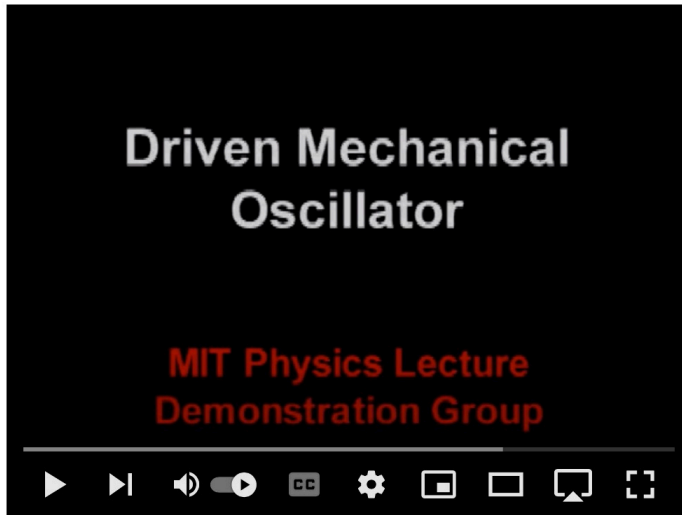
$\omega \gg \omega_0$  ➔  $\vec{p}$  Lags  $\vec{E}$  by  $\pi$

Best to stick with complex notation !

# Light-Matter Interaction

## Video of driven – damped harmonic oscillator

<https://www.youtube.com/watch?v=aZNnwQ8HJHU>



## Complex polarizability:

$$\vec{p} = e\vec{x} = e\vec{a}e^{-i(\omega t - k z)} \equiv \alpha(\omega)\vec{\epsilon}\vec{E}_0e^{-i(\omega t - k z)}$$

$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} = \frac{e^2}{m} \frac{\omega_0^2 - \omega^2 + 2i\beta\omega}{(\omega_0^2 - \omega^2)^2 + 4\beta^2\omega^2}$$

Easy to show that if  $\vec{E}(\vec{R}, t) = \vec{\epsilon}\vec{E}_0e^{-i(\omega t - k z)}$   
and  $\vec{P} = N\vec{p}$

then the wave equation reduces to

$$\left(-k^2 + \frac{\omega^2}{c^2}\right)\vec{\epsilon}\vec{E}_0e^{-i(\omega t - k z)} = -\frac{\omega^2}{c^2} \frac{N\alpha(\omega)}{\epsilon_0}\vec{\epsilon}\vec{E}_0e^{-i(\omega t - k z)}$$

→ plane wave solutions with  $k = n(\omega)\omega/c$   
where

$$k^2 = \frac{\omega^2}{c^2} \left[ 1 + \frac{N\alpha(\omega)}{\epsilon_0} \right] \equiv \frac{\omega^2}{c^2} n(\omega)$$

Complex index of refraction

# Light-Matter Interaction

Easy to show that if  $\vec{E}(\vec{R}, t) = \vec{\Sigma} E_0 e^{-i(\omega t - k z)}$   
and  $\vec{P} = N \vec{p}$

then the wave equation reduces to

$$\left(-k^2 + \frac{\omega^2}{c^2}\right) \vec{\Sigma} E_0 e^{-i(\omega t - k z)} = -\frac{\omega^2}{c^2} \frac{N \alpha(\omega)}{\epsilon_0} \vec{\Sigma} E_0 e^{-i(\omega t - k z)}$$

→ plane wave solutions with  $k = n(\omega) \omega / c$   
where

$$k^2 = \frac{\omega^2}{c^2} \left[ 1 + \frac{N \alpha(\omega)}{\epsilon_0} \right] \equiv \frac{\omega^2}{c^2} n(\omega)$$

Complex index of refraction

## Complex Index of Refraction – Physical discussion

Let

$$n(\omega) = n_R(\omega) + i n_I(\omega)$$

Plane wave propagation

$$k = n(\omega) \omega / c$$

$$\begin{aligned} \vec{E}(z, t) &= \vec{\Sigma} E_0 e^{-i(\omega t - k z)} \\ &= \vec{\Sigma} E_0 e^{-i(\omega t - [n(\omega) \omega / c] z)} \\ &= \vec{\Sigma} E_0 e^{-n_I(\omega) \omega z / c} e^{-i\omega [t - n_R(\omega) z / c]} \end{aligned}$$

We can now identify

$$\frac{c}{\omega n_I(\omega)} \leftarrow \text{attenuation length}$$

$$\frac{c}{n_R(\omega)} \leftarrow \text{phase velocity}$$

# Light-Matter Interaction

## Complex Index of Refraction – Physical discussion

Let

$$n(\omega) = n_r(\omega) + i n_i(\omega)$$

Plane wave propagation

$$k = n(\omega)\omega/c$$

$$\begin{aligned}\vec{E}(z,t) &= \vec{E}_0 e^{-i(\omega t - kz)} \\ &= \vec{E}_0 e^{-i(\omega t - [n(\omega)\omega/c]z)} \\ &= \vec{E}_0 e^{-n_i(\omega)\omega z/c} e^{-i\omega[t - n_r(\omega)z/c]}\end{aligned}$$

We can now identify

$$\frac{c}{\omega n_i(\omega)}$$



attenuation length

$$\frac{c}{n_r(\omega)}$$



phase velocity

## Absorption

The intensity of a plane wave field  $E$  is

$$\begin{aligned}I_\omega(z) &= \frac{1}{2} n_r(\omega) c \epsilon_0 |E(c,z)|^2 = I_0(0) e^{-2n_i(\omega)\omega z/c} \\ &\equiv I_0 e^{-a(\omega)z}\end{aligned}$$

where the absorption coefficient is

$$a(\omega) \equiv 2n_i(\omega)\omega/c = \frac{2\omega}{c} \text{Im} \left[ \left( 1 + \frac{N\alpha(\omega)}{\epsilon_0} \right)^{1/2} \right]$$

Possibility of gain?

No – there is no energy source!

# Light-Matter Interaction

## Absorption and Dispersion in Gases

Approximations:

$|\omega_0 - \omega| \ll \omega_0, \omega$     near resonance  
 $|n(\omega)| \sim 1$     weakly polarizable

Let  $\omega_0^2 - \omega^2 = (\omega_0 + \omega)(\omega_0 - \omega) \approx 2\omega(\omega_0 - \omega)$



$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} = \frac{e^2/2m\omega}{\omega_0 - \omega - i\beta}$$

$$= \frac{e^2}{2m\omega} \frac{\omega_0 - \omega + i\beta}{(\omega_0 - \omega)^2 + \beta^2}$$

Furthermore

$$n(\omega)^2 = 1 + \frac{N\alpha(\omega)}{\epsilon_0} = 1 + \epsilon, \epsilon \ll 1$$

Expand to 1<sup>st</sup> order  $(1 + \epsilon)^{1/2} \approx 1 + \epsilon/2$

Putting it together

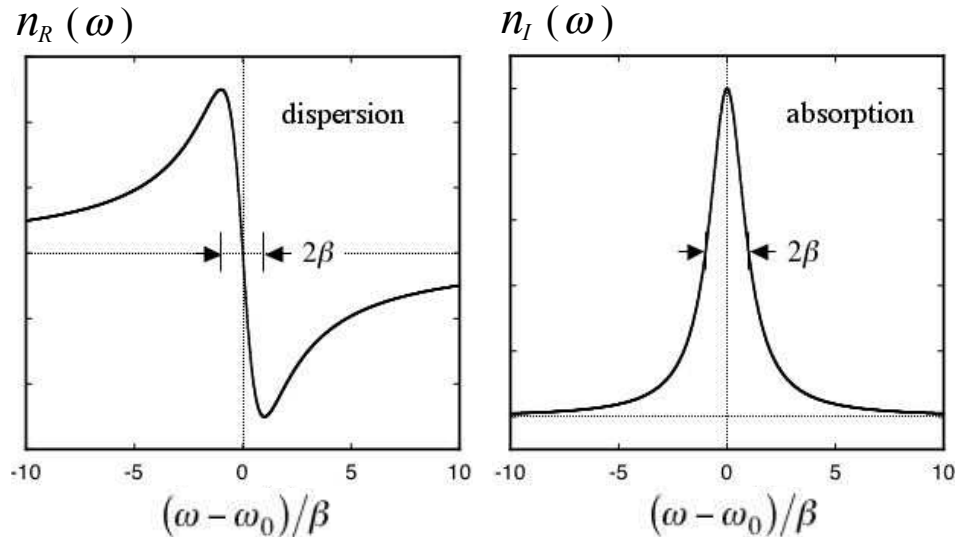
$$n_R(\omega) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\omega_0 - \omega}{(\omega_0 - \omega)^2 + \beta^2}$$

dispersive line shape

$$n_I(\omega) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{(\omega_0 - \omega)^2 + \beta^2}$$

Lorentzian line shape

General behavior:



# Light-Matter Interaction

## Putting it together

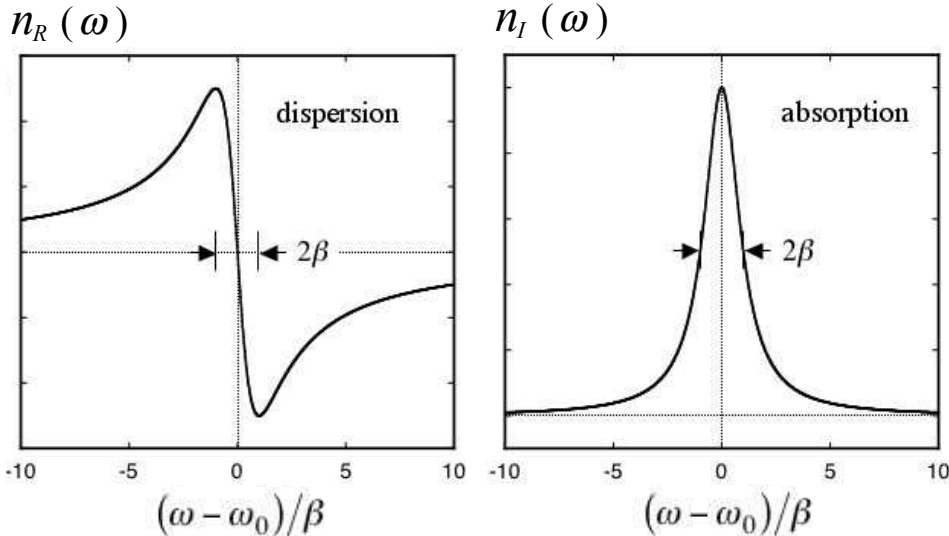
$$n_R(\omega) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\omega_0 - \omega}{(\omega_0 - \omega)^2 + \beta^2}$$

dispersive line shape

$$n_I(\omega) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{(\omega_0 - \omega)^2 + \beta^2}$$

Lorentzian line shape

## General behavior:



**Note:** for  $|\omega_0 - \omega| \gg \beta$

dispersion  $\propto \frac{1}{(\omega_0 - \omega)}$

absorption  $\propto \frac{1}{(\omega_0 - \omega)^2}$

➔ We can have loss-less dispersive media

**Note:** If we introduce the detuning  $\Delta = (\omega_0 - \omega)$  we can rewrite  $n_R(\omega)$ ,  $n_I(\omega)$  as

$$n_R(\Delta) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\Delta}{\Delta^2 + \beta^2}$$

$$n_I(\Delta) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{\Delta^2 + \beta^2}$$

From the above we see that

$$n_R(\omega) < 1 \text{ for } \omega > \omega_0 \quad \Rightarrow \quad \frac{c}{n_R(\omega)} \gg c$$

Superluminal propagation?

# Light-Matter Interaction

dispersion  $\propto \frac{1}{(\omega_0 - \omega)}$  for  $|\omega_0 - \omega| \gg \beta$

Note: absorption  $\propto \frac{1}{(\omega_0 - \omega)^2}$

➔ We can have loss-less dispersive media

Note: If we introduce the detuning  $\Delta = (\omega_0 - \omega)$  we can rewrite  $n_R(\omega)$ ,  $n_I(\omega)$  as

$$n_R(\Delta) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\Delta}{\Delta^2 + \beta^2}$$

$$n_I(\Delta) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{\Delta^2 + \beta^2}$$

From the above we see that

$$n_R(\omega) < 1 \text{ for } \omega > \omega_0 \quad \Rightarrow \quad \frac{c}{n_R(\omega)} \gg c$$

Superluminal propagation?

## Free Electrons

Consider the limit  $\omega \gg \omega_0$

➔ effectively unbound electrons

This is a reasonable model of plasmas & metals

In this limit we have

$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} \approx -\frac{e^2}{m\omega} \quad \Rightarrow$$

$$n(\omega) = \sqrt{1 + \frac{N(\alpha)}{\epsilon_0}} \approx \sqrt{1 - \frac{Ne^2}{\epsilon_0 m \omega^2}} \equiv \sqrt{1 - \frac{\omega_p^2}{\omega^2}}$$

We introduce the Plasma Frequency :

# Light-Matter Interaction

dispersion  $\propto \frac{1}{(\omega_0 - \omega)}$  for  $|\omega_0 - \omega| \gg \beta$

Note: absorption  $\propto \frac{1}{(\omega_0 - \omega)^2}$

➔ We can have loss-less dispersive media

Note: If we introduce the detuning  $\Delta = (\omega_0 - \omega)$  we can rewrite  $n_R(\omega)$ ,  $n_I(\omega)$  as

$$n_R(\Delta) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\Delta}{\Delta^2 + \beta^2}$$

$$n_I(\Delta) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{\Delta^2 + \beta^2}$$

From the above we see that

$$n_R(\omega) < 1 \text{ for } \omega > \omega_0 \Rightarrow \frac{c}{n_R(\omega)} \gg c$$

Superluminal propagation?

## Free Electrons

Consider the limit  $\omega \gg \omega_0$

➔ effectively unbound electrons

This is a reasonable model of plasmas & metals

In this limit we have

$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} \approx -\frac{e^2}{m\omega}$$

$$n(\omega) = \sqrt{1 + \frac{N(\alpha)}{\epsilon_0}} \approx \sqrt{1 - \frac{Ne^2}{\epsilon_0 m \omega^2}} \equiv \sqrt{1 - \frac{\omega_p^2}{\omega^2}}$$

We introduce the Plasma Frequency :

$$\omega_p = \sqrt{\frac{Ne^2}{\epsilon_0 m}}$$



# Light-Matter Interaction

dispersion  $\propto \frac{1}{(\omega_0 - \omega)}$  for  $|\omega_0 - \omega| \gg \beta$

Note: absorption  $\propto \frac{1}{(\omega_0 - \omega)^2}$

➔ We can have loss-less dispersive media

Note: If we introduce the detuning  $\Delta = (\omega_0 - \omega)$  we can rewrite  $n_R(\omega)$ ,  $n_I(\omega)$  as

$$n_R(\Delta) = 1 + \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\Delta}{\Delta^2 + \beta^2}$$

$$n_I(\Delta) = \frac{Ne^2}{4\epsilon_0 m \omega} \frac{\beta}{\Delta^2 + \beta^2}$$

From the above we see that

$$n_R(\omega) < 1 \text{ for } \omega > \omega_0 \Rightarrow \frac{c}{n_R(\omega)} \gg c$$

Superluminal propagation?

## Free Electrons

Consider the limit  $\omega \gg \omega_0$

➔ effectively unbound electrons

This is a reasonable model of plasmas & metals

In this limit we have

$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} \approx -\frac{e^2}{m\omega}$$

$$n(\omega) = \sqrt{1 + \frac{N(\alpha)}{\epsilon_0}} \approx \sqrt{1 - \frac{Ne^2}{\epsilon_0 m \omega^2}} \equiv \sqrt{1 - \frac{\omega_p^2}{\omega^2}}$$

We introduce the Plasma Frequency :

$$\omega_p = \sqrt{\frac{Ne^2}{\epsilon_0 m}}$$

# Light-Matter Interaction

## Free Electrons

Consider the limit  $\omega \gg \omega_0$

➔ effectively unbound electrons

This is a reasonable model of plasmas & metals

In this limit we have

$$\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 - 2i\beta\omega} \approx -\frac{e^2}{m\omega} \quad \rightarrow$$

$$n(\omega) = \sqrt{1 + \frac{N(\alpha)}{\epsilon_0}} \approx \sqrt{1 - \frac{Ne^2}{\epsilon_0 m \omega^2}} \equiv \sqrt{1 - \frac{\omega_p^2}{\omega^2}}$$

We introduce the Plasma Frequency :

$$\omega_p = \sqrt{\frac{Ne^2}{\epsilon_0 m}}$$

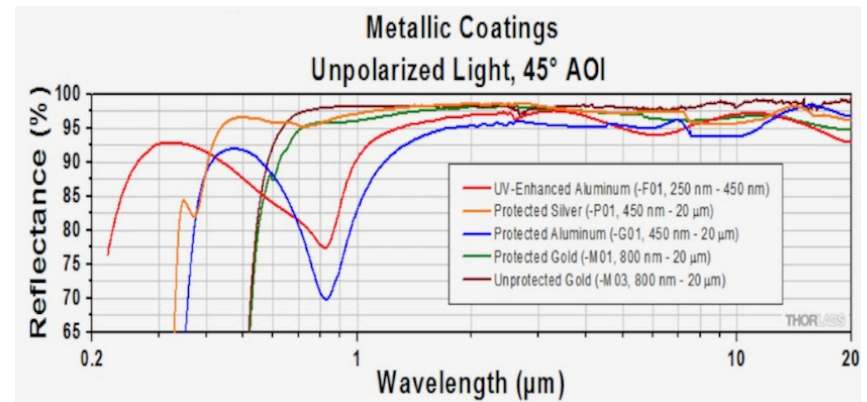
Let  $\left. \begin{array}{l} \omega_0 \ll \omega \ll \omega_p \\ |\omega_0 - \omega| \gg \beta \end{array} \right\} \rightarrow n(\omega) \text{ purely imaginary} \\ \text{- but no loss!}$

We now have

$$\begin{aligned} \vec{E}(z,t) &= \vec{\epsilon} E_0 e^{-i\omega[t - n(\omega)z/c]} \\ &= \vec{\epsilon} E_0 e^{-i\omega t} e^{i(z/c)\sqrt{\omega^2 - \omega_p^2}} \\ &= \vec{\epsilon} E_0 e^{-i\omega t} e^{-b(\omega)z} \end{aligned}$$

where  $b(\omega) = -\frac{i}{c} \sqrt{\omega^2 - \omega_p^2}$

Reflection at surface, penetration depth  $\sim 1/b(\omega)$



Examples of this kind of medium includes plasmas, and metals such as aluminum, silver and gold which are known to be excellent mirrors for visible and IR radiation.

