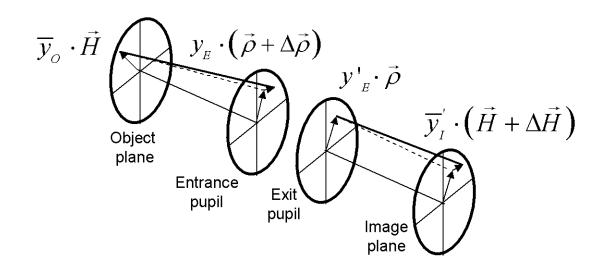
Introduction to aberrations

OPTI 518 Lectures 9





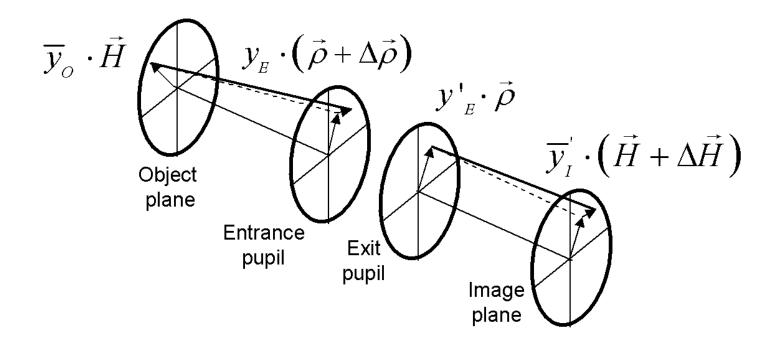
Topics

- Transverse ray aberrations
- RMS spot size
- Spot diagrams
- Ray fans
- Special locations along the spherical aberration caustic



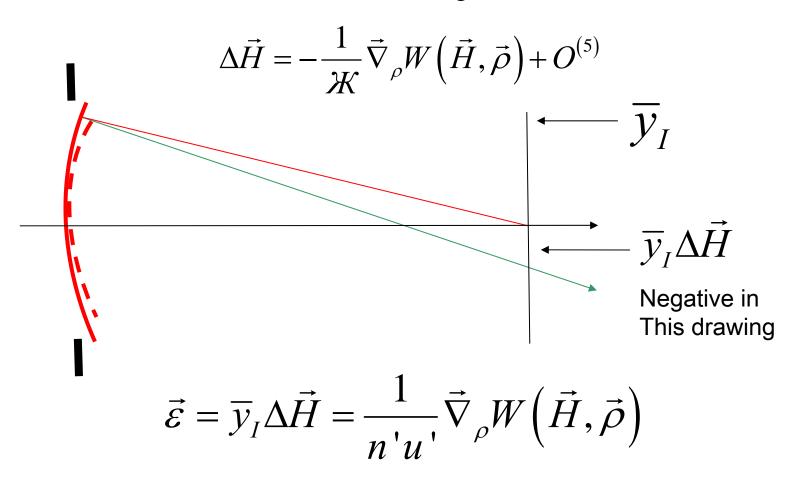
Transverse ray aberration

$$\Delta \vec{H} = -\frac{1}{\mathcal{K}} \vec{\nabla}_{\rho} W \left(\vec{H}, \vec{\rho} \right) + O^{(5)}$$





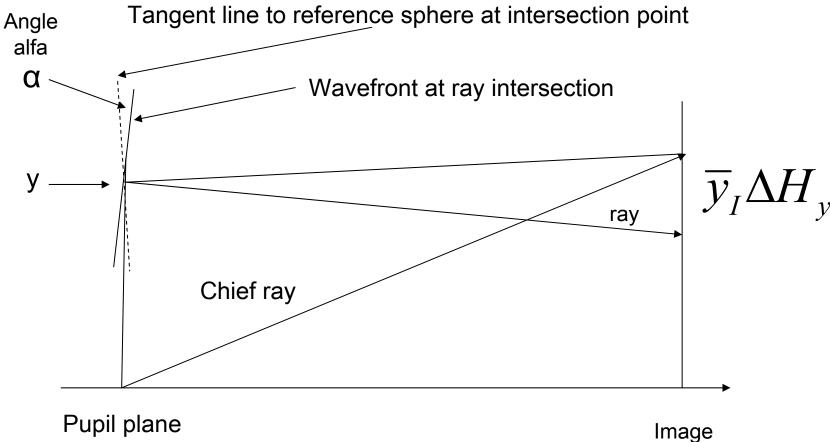
Transverse ray aberration



 $\overline{\mathcal{Y}}_I$ Is the chief ray height at the image plane



Derivation

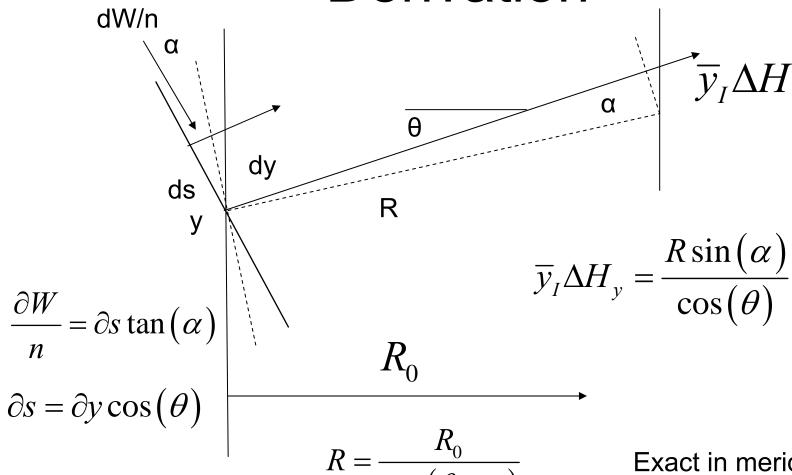


We exactly know where the real ray intersects the exit pupil plane as we have selected it with the tip of the aperture vector

lmage plane



Derivation



Exact in meridional plane



Derivation

$$\frac{1}{n} \frac{\partial W}{\partial y} = \frac{\cos^3(\theta)}{R_0} \left(1 - \tan(\alpha) \tan(\theta) \right) \overline{y}_I \Delta H_y$$

$$= \frac{\cos^3(\theta)}{R_0} \overline{y}_I \Delta H_y + O^{(7)} = \frac{1}{R_0} \overline{y}_I \Delta H_y + O^{(5)}$$



Derivation to 2D

$$\frac{1}{n} \frac{y}{u} \frac{\partial W}{\partial y} = \frac{1}{nu} \frac{\partial W}{\partial \rho_{y}} = \overline{y}_{I} \Delta H_{y} + O^{(5)}$$

$$\frac{-1}{\mathcal{K}} \frac{\partial W}{\partial \rho_{y}} = \Delta H_{y} + O^{(5)}$$

$$\frac{-1}{\mathcal{K}} \frac{\partial W}{\partial \rho_{x}} = \Delta H_{x} + O^{(5)}$$

$$\frac{-1}{\mathcal{K}} \vec{\nabla}_{\rho} W = \Delta \vec{H} + O^{(5)}$$

Key relationship between wave and transverse aberrations



Example

$$\vec{\varepsilon} = \overline{y}_{I} \Delta \vec{H} = \frac{1}{n'u'} \vec{\nabla}_{\rho} W (\vec{H}, \vec{\rho})$$

$$= \frac{1}{n'u'} \vec{\nabla}_{\rho} W_{040} (\vec{\rho} \cdot \vec{\rho})^{2}$$

$$= \frac{4}{n'u'} W_{040} (\vec{\rho} \cdot \vec{\rho}) \vec{\rho}$$

$$= \frac{4}{n'u'} W_{040} \rho^{2} (\rho_{h} \vec{h} + \rho_{i} \vec{i})$$



Gradient operator

$$\vec{\nabla}_{\rho} (\rho^{2}) = \vec{\nabla}_{\rho} (\rho_{h}^{2} + \rho_{i}^{2}) = \frac{\partial (\rho_{h}^{2} + \rho_{i}^{2})}{\partial \rho_{h}} \vec{h} + \frac{\partial (\rho_{h}^{2} + \rho_{i}^{2})}{\partial \rho_{i}} \vec{i}$$

$$= 2(\rho_{h} \vec{h} + \rho_{i} \vec{i}) = 2\vec{\rho}$$

$$\vec{\nabla}_{\rho} \left(\vec{\rho} \cdot \vec{\rho} \right) = 2\vec{\rho}$$

$$\vec{\nabla}_{\rho} \left(\vec{H} \cdot \vec{\rho} \right) = \vec{H}$$

$$\vec{H} \cdot \vec{\rho} = H_h \rho_h + H_i \rho_i$$



Normalized mean square spot size and RMS spot size

$$\left| \overline{\Delta \vec{H}} \right|^{2} = \frac{1}{\mathcal{K}^{2}} \left| \overline{\nabla}_{\rho} W \left(\vec{H}, \vec{\rho} \right) \right|^{2} = \frac{1}{\mathcal{K}^{2}} \frac{1}{\pi} \int_{0}^{2\pi} \int_{0}^{1} \left| \overline{\nabla}_{\rho} W \left(\vec{H}, \vec{\rho} \right) \right|^{2} \rho d\rho d\phi$$

$$RMS = \left| \overline{y}_I \right| \sqrt{\left| \Delta \vec{H} \right|^2}$$

RMS spot size is an image quality criteria



Spherical aberration and defocus

$$\overline{y}_{I} \Delta \vec{H} = \frac{1}{n'u'} \vec{\nabla}_{\rho} W (\vec{H}, \vec{\rho})$$

$$= \frac{1}{n'u'} \vec{\nabla}_{\rho} \left(W_{040} (\vec{\rho} \cdot \vec{\rho})^{2} + W_{020} (\vec{\rho} \cdot \vec{\rho}) \right)$$

$$= \frac{1}{n'u'} \left(4W_{040} (\vec{\rho} \cdot \vec{\rho}) \vec{\rho} + 2W_{020} \vec{\rho} \right)$$

$$\Delta \vec{H} \cdot \Delta \vec{H} = \frac{1}{\mathcal{K}^{2}} \left| \vec{\nabla}_{\rho} W \left(\vec{H}, \vec{\rho} \right) \right|^{2} = \frac{1}{\mathcal{K}^{2}} \left| \left(4W_{040} \left(\vec{\rho} \cdot \vec{\rho} \right) \vec{\rho} + 2W_{020} \vec{\rho} \right) \right|^{2}$$

$$= \frac{1}{\mathcal{K}^{2}} \left(16W_{040}^{2} \left(\vec{\rho} \cdot \vec{\rho} \right)^{3} + 16W_{040} W_{020} \left(\vec{\rho} \cdot \vec{\rho} \right)^{4} + 4W_{020}^{2} \left(\vec{\rho} \cdot \vec{\rho} \right) \right)$$



Spherical aberration and defocus

$$\begin{split} & \overline{\left|\Delta \vec{H}\right|^2} = \frac{1}{\mathcal{K}^2} \frac{1}{\pi} \int_{0}^{2\pi} \int_{0}^{1} \left(16W_{040}^2 \rho^6 + 16W_{040}W_{020}\rho^4 + 4W_{020}^2 \rho^2\right) \rho d\rho d\phi \\ & = \frac{2}{\mathcal{K}^2} \left(\frac{16}{8}W_{040}^2 + \frac{16}{6}W_{040}W_{020} + \frac{4}{4}W_{020}^2\right) \\ & = \frac{1}{\mathcal{K}^2} \left(4W_{040}^2 + \frac{16}{3}W_{040}W_{020} + 2W_{020}^2\right) \\ & = \frac{1}{\mathcal{K}^2} \left(2\left(W_{020} + \frac{4}{3}W_{040}\right)^2 + \frac{4}{9}W_{040}^2\right) \end{split}$$



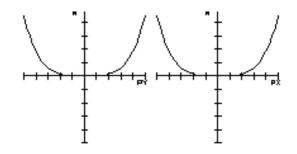
"Best focus"

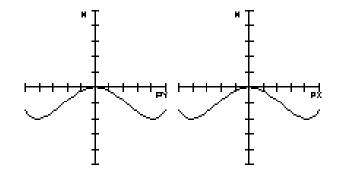
Minimum rms spot when:

$$W_{020} = -\frac{4}{3}W_{040}$$

Normalized RMS spot size

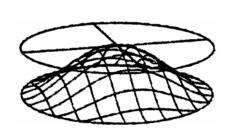
$$\sqrt{\left|\Delta \vec{H}\right|^2} = \frac{1}{\mathcal{K}} \frac{2}{3} W_{040}$$

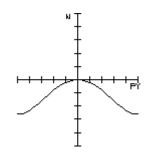






Marginal focus





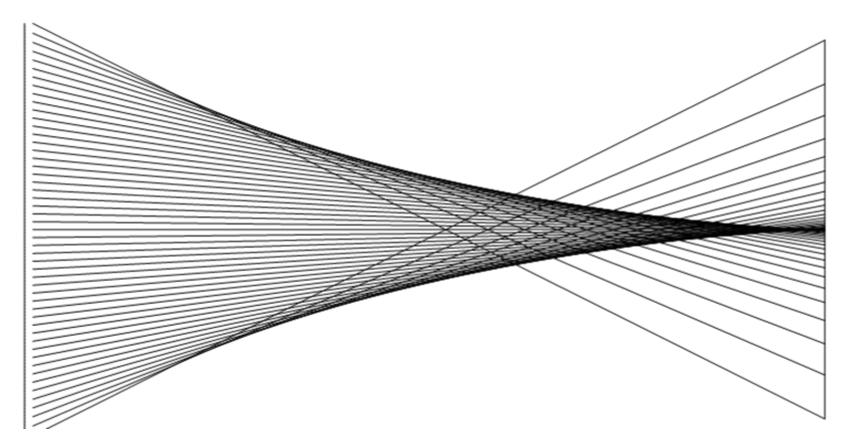
$$\begin{split} & \overline{y}_{I} \Delta \vec{H} = \frac{1}{n'u'} \vec{\nabla}_{\rho} W (\vec{H}, \vec{\rho}) \\ & = \frac{1}{n'u'} \vec{\nabla}_{\rho} \left(W_{040} (\vec{\rho} \cdot \vec{\rho})^{2} + W_{020} (\vec{\rho} \cdot \vec{\rho}) \right) \\ & = \frac{1}{n'u'} \left(4W_{040} (\vec{\rho} \cdot \vec{\rho}) \vec{\rho} + 2W_{020} \vec{\rho} \right) \end{split}$$

Marginal focus is at:

$$W_{020} = -2W_{040}$$



Ray caustic for spherical aberration



Caustic ~ burning



Special locations along the ray caustic

$$W_{020} = -W_{040}$$

Minimum wavefront variance

$$W_{020} = -\frac{4}{3}W_{040}$$

Minimum rms spot size

$$W_{020} = -\frac{3}{2}W_{040}$$

Minimum circle



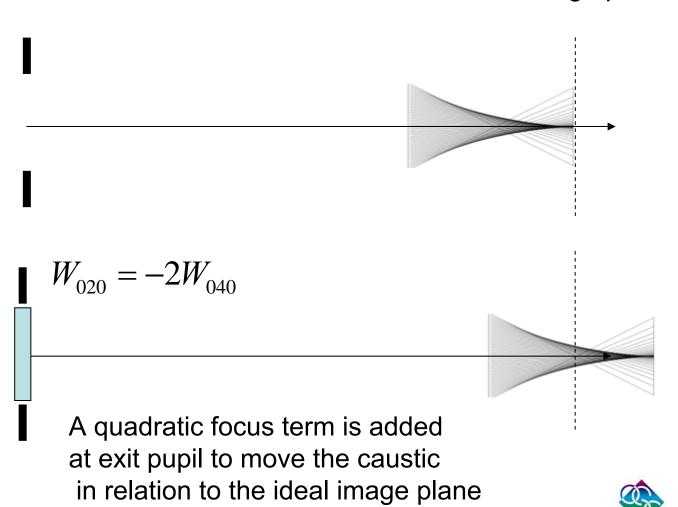
$$W_{020} = -2W_{040}$$

Marginal focus



Focus view

Ideal image plane



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Wave deformation shapes

Wave aberration shapes						
Zero-order						
W_{000}						
Second-order						
$W_{020}(\vec{ ho}\cdot\vec{ ho})$	$W_{111}ig(ec{H}\cdotec{ ho}ig)$	$W_{200}ig(ec{H}\cdotec{H}ig)$				
Fourth-order						
$W_{040}(\vec{ ho}\cdot\vec{ ho})^2$	$W_{131}(\vec{H}\cdot\vec{ ho})(\vec{ ho}\cdot\vec{ ho})$	$W_{222}ig(ec{H}\cdotec{ ho}ig)^2$				
$W_{220}(\vec{H}\cdot\vec{H})(\vec{\rho}\cdot\vec{\rho})$	$W_{311}(\vec{H}\cdot\vec{H})(\vec{H}\cdot\vec{\rho})$	$W_{400}ig(ec{H}\cdotec{H}ig)^2$				

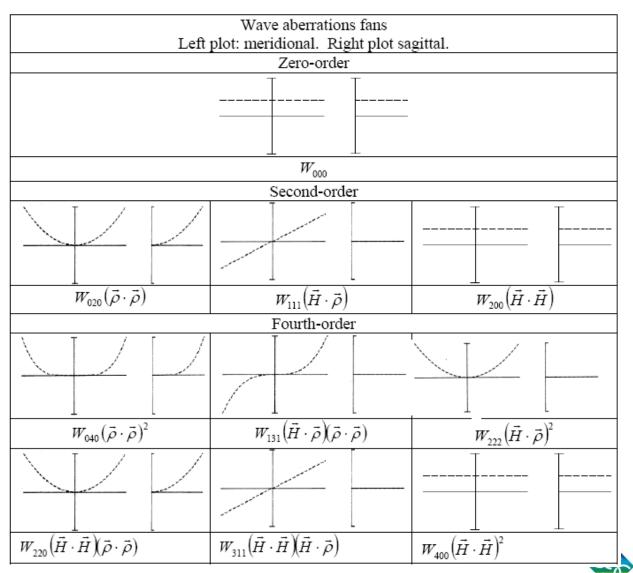


Notes

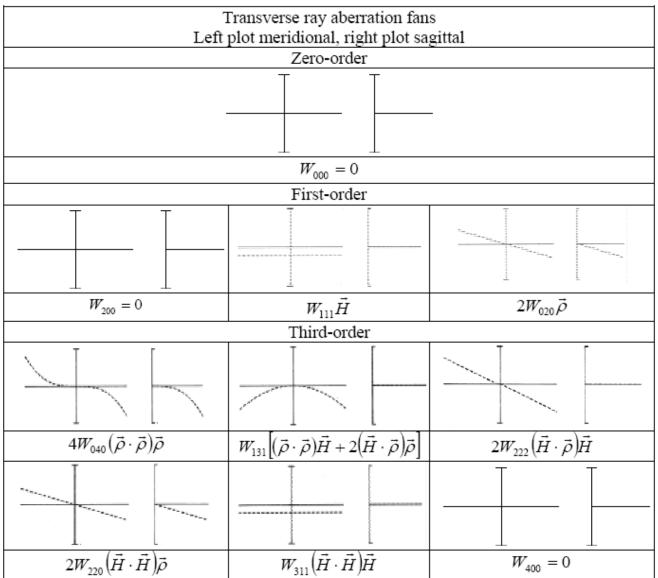
- Pure terms
- Fourth-order terms
- Actually they are a mixture including higher order.
- Need to learn them to recognize dominant aberrations
- Preferred in waves



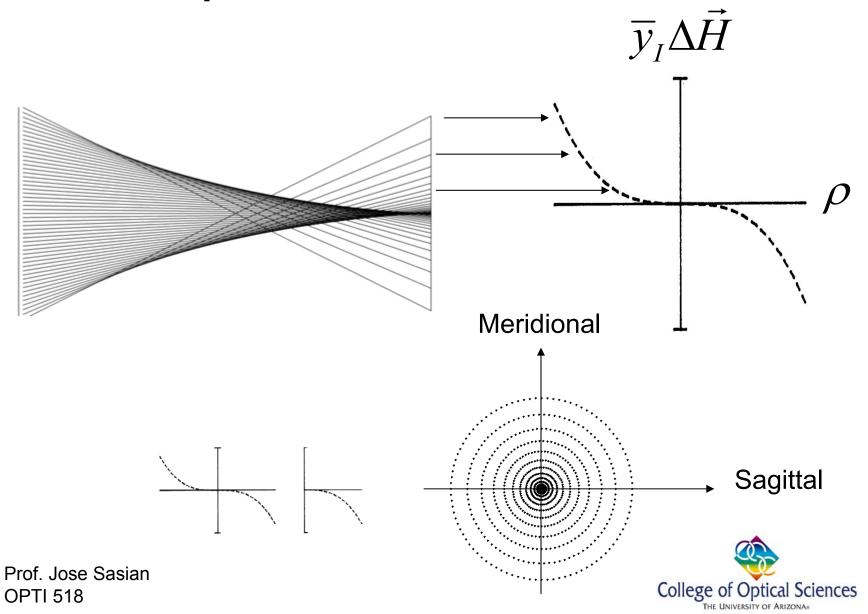
Wave fans



Ray fans



Spherical aberration



Transverse ray aberrations

$$\vec{\nabla}_{\rho}W = \sum_{j,m,n} W_{k,l,m} (\vec{H} \cdot \vec{H})^{j} \left[(\vec{\rho} \cdot \vec{\rho})^{n} \vec{\nabla}_{\rho} (\vec{H} \cdot \vec{\rho})^{m} + (\vec{H} \cdot \vec{\rho})^{m} \vec{\nabla}_{\rho} (\vec{\rho} \cdot \vec{\rho})^{n} \right] =$$

$$= \sum_{j,m,n} W_{k,l,m} (\vec{H} \cdot \vec{H})^{j} \left[m (\vec{H} \cdot \vec{\rho})^{m-1} (\vec{\rho} \cdot \vec{\rho})^{n} \vec{\nabla}_{\rho} (\vec{H} \cdot \vec{\rho}) + n (\vec{H} \cdot \vec{\rho})^{m} (\vec{\rho} \cdot \vec{\rho})^{n-1} \vec{\nabla}_{\rho} (\vec{\rho} \cdot \vec{\rho}) \right] =$$

$$= \sum_{j,m,n} W_{k,l,m} (\vec{H} \cdot \vec{H})^{j} \left[m (\vec{H} \cdot \vec{\rho})^{m-1} (\vec{\rho} \cdot \vec{\rho})^{n} \vec{H} + 2n (\vec{H} \cdot \vec{\rho})^{m} (\vec{\rho} \cdot \vec{\rho})^{n-1} \vec{\rho} \right]$$

The gradient of the wavefront has two component vectors; one in the direction of the field vector and another in the direction of the aperture vector. Note that the transverse ray aberrations have one algebraic order less than the wavefront aberrations.



Transverse ray aberrations

Transverse ray aberrations				
Aberration	Vector form			
name/order				
Zero order				
Uniform piston	0			
First order Gaussian				
Quadratic piston	0			
Magnification	$W_{111}ec{H}$			
Focus	$2W_{020}\vec{\rho}$			
Third order Seidel				
Spherical aberration	$4W_{040}(\vec{ ho}\cdot\vec{ ho})\vec{ ho}$			
Coma	$W_{131}\left[(\vec{\rho} \cdot \vec{\rho}) \vec{H} + 2 \left(\vec{H} \cdot \vec{\rho} \right) \vec{\rho} \right]$			
Astigmatism	$2W_{222}(\vec{H}\cdot\vec{\rho})\vec{H}$			
Field curvature	$2W_{220}(\vec{H}\cdot\vec{H})\vec{\rho}$			
Distortion	$W_{311}(\vec{H}\cdot\vec{H})\vec{H}$			
Quartic piston	0			

Piston terms do not contribute transverse ray aberration



Components

In order to calculate and plot the components of the transverse ray aberrations we write the field and aperture vectors as,

$$\vec{H} = H\vec{h}$$
$$\vec{\rho} = \rho \vec{g}$$

where \vec{h} is a unit vector in the direction of \vec{H} and \vec{g} is a unit vector in the direction of $\vec{\rho}$. The magnitudes H and ρ have ranges $0 \le H \le 1$ and $0 \le \rho \le 1$. The transverse ray aberration vector $\vec{y}_I \Delta \vec{H}$ can be written as,

$$\overline{y}_I \Delta \vec{H} = A\vec{h} + B\vec{g}$$



Components

$$A = \frac{1}{n'u'} \sum_{j,m,m} W_{k,l,m} \left[m (\vec{H} \cdot \vec{H})^{j} (\vec{H} \cdot \vec{\rho})^{m-1} (\vec{\rho} \cdot \vec{\rho})^{n} \right] H =$$

$$= \frac{1}{n'u'} \sum_{j,m,m} W_{k,l,m} H^{k} \rho^{l-1} \left[m \cos^{m-1} (\phi) \right]$$

$$B = \frac{1}{n'u'} \sum_{j,m,m} W_{k,l,m} \left[2n \left(\vec{H} \cdot \vec{H} \right)^j \left(\vec{H} \cdot \vec{\rho} \right)^m \left(\vec{\rho} \cdot \vec{\rho} \right)^{n-1} \right] \rho =$$

$$= \frac{1}{n'u'} \sum_{l,m,m} W_{k,l,m} H^{k} \rho^{l-1} [(l-m)\cos^{m}(\phi)]$$

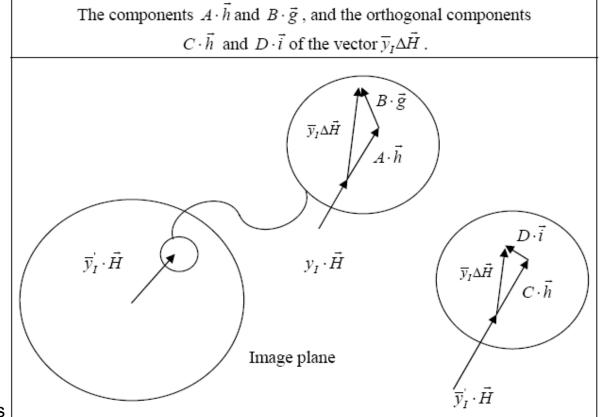
$$k = 2j + m \qquad l = 2n + m$$



The components of $\overline{y}_I \Delta H$

$$\overline{y}_{I}\Delta\vec{H} = A\vec{h} + B\vec{g}$$

$$\overline{y}_{I}\Delta\vec{H} = C\vec{h} + D\vec{i} = (A + B\cos(\varphi))\vec{h} + (B\sin(\varphi))\vec{i}$$



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Orthogonal Components

$$\overline{y}_{I} \Delta \vec{H} = C\vec{h} + D\vec{i} = (A + B\cos(\varphi))\vec{h} + (B\sin(\varphi))\vec{i}$$

 $ec{i}$ is a unit vector orthogonal to $ec{h}$

$$C = \frac{1}{n'u'} \sum_{i,m,m} W_{k,l,m} H^{k} \rho^{l-1} \left[m \cos^{m-1}(\phi) + (l-m) \cos^{m}(\phi) \cos(\phi) \right]$$

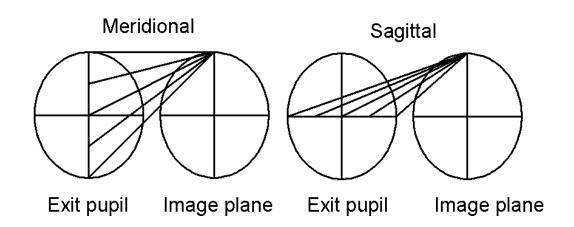
$$D = \frac{1}{n'u'} \sum_{l,m,m} W_{k,l,m} H^{k} \rho^{l-1} [(l-m)\cos^{m}(\phi)\sin(\phi)]$$



$$k = 2j + m \qquad l = 2n + m$$

Choice of ray fan plots

For historical reasons plots of ray intercepts at the observation plane, of rays in the meridional plane and in the sagittal plane, are usually produced. Since there is symmetry only one half of the sagittal rays intercepts may be plotted.



$$(\phi = 0) \qquad (\phi = 90)$$



Meridional rays

$$C = \frac{1}{n'u'} \sum_{j,m,m} W_{k,l,m} H^{k} \rho^{l-1} \Big[m \cos^{m-1}(\phi) + (l-m) \cos^{m}(\phi) \cos(\phi) \Big]$$

$$D = \frac{1}{n'u'} \sum_{j,m,m} W_{k,l,m} H^{k} \rho^{l-1} \Big[(l-m) \cos^{m}(\phi) \sin(\phi) \Big]$$

For the meridional rays the intercept is given by, $(\phi = 0)$

$$C = \frac{1}{n'u'} \sum_{j,m,n} W_{k,l,m} H^{k} (l\rho^{l-1})$$

$$D = 0$$

$$l = 2n + m$$

$$k = 2j + m$$

Since D = 0

the meridional ray intercepts remain in the meridional plane.

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Sagittal rays

$$(\phi = 90)$$

According to the integer m the sagittal ray intercepts are given by,

$$C = \frac{1}{n'u'} \sum_{j,1,n} W_{k,l,1} H^{k} (\rho^{l-1})$$

$$m = 1$$

$$C = 0$$

$$m \neq 1$$

$$D = \frac{1}{n'u'} \sum_{j,0,n} W_{k,l,0} H^{k} (l\rho^{l-1})$$

$$m = 0$$

$$D = 0$$

$$m \neq 0$$

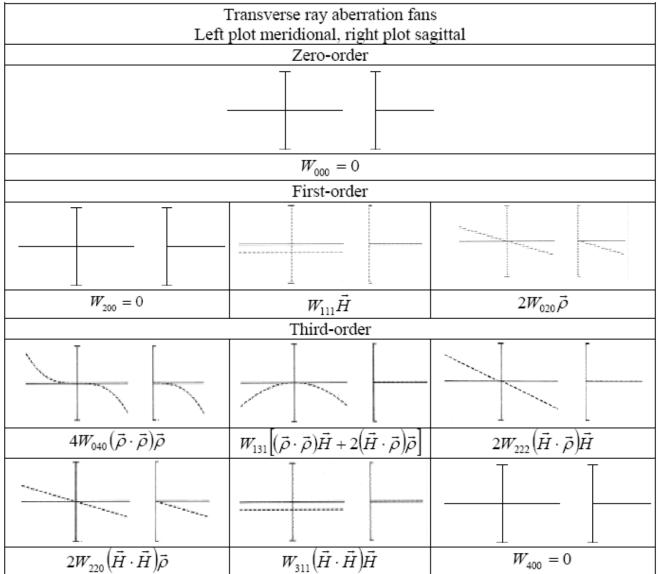


Summary

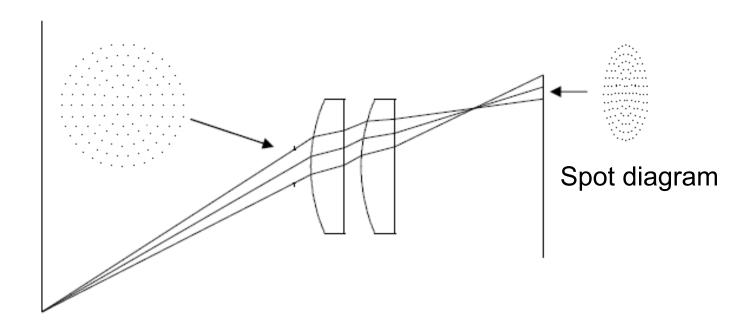
Transverse ray aberrations scalar components					
Aberration	A	В	С	D	
W_{000}	0	0	0	0	
$W_{200}ig(ec{H}\cdotec{H}ig)$	0	0	0	0	
$W_{111}(\vec{H}\cdot\vec{ ho})$	Н	0	Н	0	
$W_{020}(\vec{ ho}\cdot\vec{ ho})$	0	2ρ	$2\rho\cos(\phi)$	$2\rho\sin(\phi)$	
$W_{040}(\vec{ ho}\cdot\vec{ ho})^2$	0	$4\rho^3$	$4\rho^3\cos(\phi)$	$4\rho^3\sin(\phi)$	
$W_{131}(\vec{H}\cdot\vec{ ho})(\vec{ ho}\cdot\vec{ ho})$	$H\rho^2$	$2H\rho^2\cos(\phi)$	$H\rho^2 (1 + 2\cos^2(\phi))$	$2H\rho^2\sin(\phi)\cos(\phi)$	
$W_{222} (\vec{H} \cdot \vec{ ho})^2$	$2H\rho^2\cos(\phi)$	0	$2H\rho^2\cos(\phi)$	0	
$W_{220} \left(\vec{H} \cdot \vec{H} \right) \left(\vec{ ho} \cdot \vec{ ho} \right)$	0	$2H\rho^2$	$2H\rho^2\cos(\phi)$	$2H\rho^2\sin(\phi)$	
$W_{311}(\vec{H}\cdot\vec{H})(\vec{H}\cdot\vec{\rho})$	H^3	0	H^3	0	
$W_{400}ig(ec{H}\cdotec{H}ig)^2$	0	0	0	0	



Transverse ray aberration fans



Concept of spot diagram

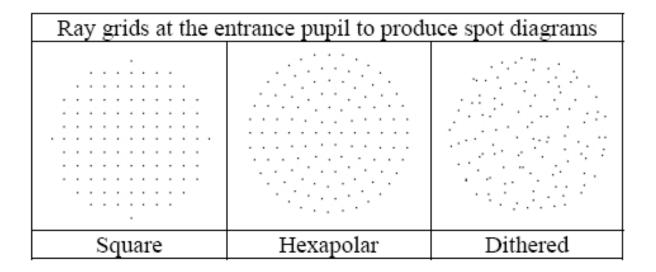


A grid of points is defined at the entrance or exit pupil of an optical system and rays are traced from the object point through the grid points, through the observation plane.

The rays intersections form a spot diagram.



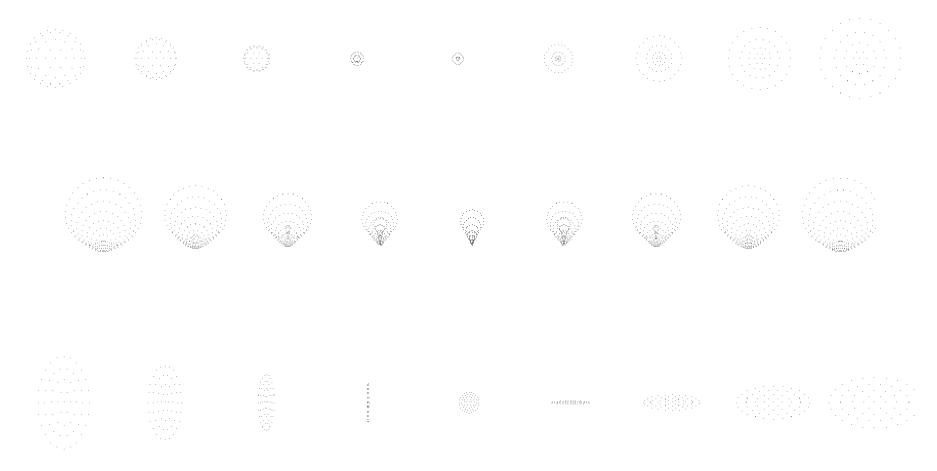
Ray grids



Located at the entrance or exit pupil plane



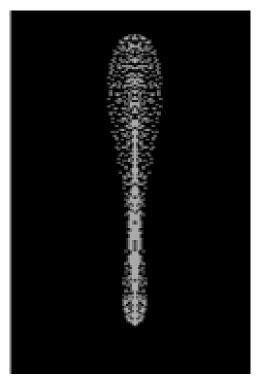
Spot diagrams for spherical aberration, coma, and astigmatism



What are the symmetries?



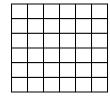
Binning and counting spots





Spot diagram

Binning and counting



Diffraction based



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Summary

- Transverse ray aberration
- Spot diagrams
- Wave fans and ray fans
- Meridional and sagittal planes
- Special locations along the ray caustic
- RMS spot size

