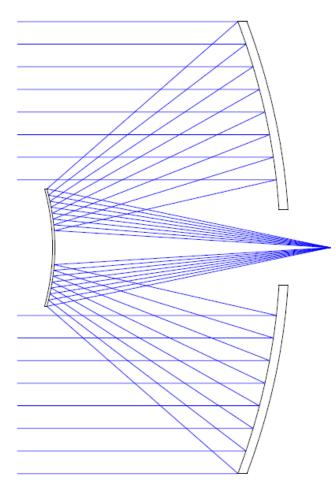
# Some Reflective Systems

# Introduction to aberrations OPTI 518



# Cassegrain type



- •True Cassegrain
- •Ritchey-Chretien: aplanatic
- •Dall-Kirkham: spherical secondary
- Pressman-Camichel; spherical primary
- Olivier Guyon (no diffraction rings)



### Structural coefficients

Seidel sums in terms of structural aberration coefficients

> Pupils located at principal planes

$$S_I = \frac{1}{4} y_P^4 \Phi^3 \sigma_I$$

$$S_{II} = \frac{1}{2} \mathcal{H} y_P^2 \Phi^2 \sigma_{II}$$

$$S_{III} = \mathcal{K}^2 \Phi \sigma_{III}$$

$$S_{IV} = \mathcal{K}^2 \Phi \sigma_{IV}$$

$$S_V = \frac{2\mathcal{K}^3 \sigma_V}{y_P^2}$$

$$C_L = y_P^2 \Phi \sigma_L$$

$$C_T = 2\mathcal{K}\sigma_T$$

Stop-shift from principal planes
$$\sigma_{I}^{*} = \sigma_{I}$$

$$\sigma_{II}^{*} = \sigma_{II} + \overline{S}_{\sigma} \sigma_{I}$$

$$\sigma_{III}^{*} = \sigma_{III} + 2\overline{S}_{\sigma} \sigma_{II} + \overline{S}_{\sigma}^{2} \sigma_{I}$$

$$\sigma_{IV}^{*} = \sigma_{IV}$$

$$\sigma_{V}^{*} = \sigma_{V} + \overline{S}_{\sigma} \left(\sigma_{IV} + 3\sigma_{III}\right) + 3\overline{S}_{\sigma}^{2} \sigma_{II} + \overline{S}_{\sigma}^{3} \sigma_{I}$$

$$\sigma_{L}^{*} = \sigma_{L}$$

$$\sigma_{T}^{*} = \sigma_{T} + \overline{S}_{\sigma} \sigma_{L}$$

$$\overline{S}_{\sigma} = \frac{y_{P} \overline{y}_{P} \Phi}{2\mathcal{K}}$$

$$\Delta \overline{S}_{\sigma} = \frac{y_{P} \Delta \overline{y}_{P} \Phi}{2\mathcal{K}} = \frac{y_{P}^{2} \Phi}{2\mathcal{K}} \overline{S}$$



Structural aberration coefficients of				
a reflecting surface in air				
Stop at surface	e With stop shift			
$\sigma_I = Y^2 + \alpha$	$\sigma_I = Y^2 + \alpha$			
$\sigma_{II} = -Y$	$\sigma_{II} = -Y \left( 1 - \overline{S}_{\sigma} Y \right) + \overline{S}_{\sigma} \cdot \alpha$			
$\sigma_{III} = 1$	$\sigma_{III} = \left(1 - \overline{S}_{\sigma}Y\right)^2 + \overline{S}_{\sigma}^2 \cdot \alpha$			
$\sigma_{IV} = -1$	$\sigma_{IV} = -1$			
$\sigma_{V} = 0$	$\sigma_{V} = \overline{S}_{\sigma} \cdot \left(1 - \overline{S}_{\sigma}Y\right) \cdot \left(2 - \overline{S}_{\sigma}Y\right) + \overline{S}_{\sigma}^{3} \cdot \alpha$			

$$\alpha = K = -\varepsilon^2$$



# Aspheric mirrors

$$\frac{CONIC}{4y_{p}^{*}\phi^{3}x = Ky^{4}c^{3}\Delta(n) = \frac{Ky^{4}\phi_{s}^{3}}{[\Delta(n)]^{2}}}$$

$$X = \frac{4K(y/y_{p})^{4}(\phi_{s}/\phi)^{3}}{[\Delta(n)]^{2}} \quad K = \frac{[\Delta(n)]^{2}}{4(y/y_{p})^{4}(\phi_{s}/\phi)^{3}}$$

EENERAL

$$4y_{P}^{*}\phi^{3}x = 8a_{4}y^{4}\Delta(n)$$
 $x = \frac{32a_{4}(4/y_{P})^{4}\Delta(n)}{\phi^{3}}$ 
 $a_{4} = \frac{\phi^{3}x}{32(4/y_{P})^{4}\Delta(n)}$ 

$$\sqrt{1/4} = M$$
 $\sqrt{2/4} = (1-M)(1+ML)$ 
 $\sqrt{4/4} = 1$ 
 $\sqrt{4^2/4} = \frac{1}{1+ML}$ 
 $\sqrt{2} = 0$ 
 $\sqrt{2} = \frac{1}{1+ML}$ 
 $\sqrt{2} = \frac{1}{1+ML}$ 
 $\sqrt{2} = \frac{1}{1+ML}$ 
 $\sqrt{2} = \frac{1+M}{1-ML}$ 

$$L = \frac{y_B}{y_B} \qquad M = \frac{1 - y_B}{y_B}$$



#### COMPONENT ABERRATIONS

$$\sigma_{\overline{n}} = 1$$



#### SYSTEM ABERRATIONS

$$\sigma_{\mathrm{I}} = M^{3}(1+\alpha_{1}) + K\left[\left(\frac{1+M}{1-M}\right)^{2} + \alpha_{2}\right]$$

$$\sigma_{\overline{m}} = -1 + \left(\frac{1}{2}\right) K \left[\left(\frac{1+M}{1-M}\right)^2 + \times_2\right]$$



SPHERICAL CORRECTION

$$\alpha_{1} = -1 \qquad \alpha_{2} = -\left(\frac{1+M}{1-M}\right)^{2}$$

$$\left[\left(\frac{1+M}{1-M}\right)^{2} + \alpha_{2}\right] = 0$$

#### APLANATIC CORRECTION

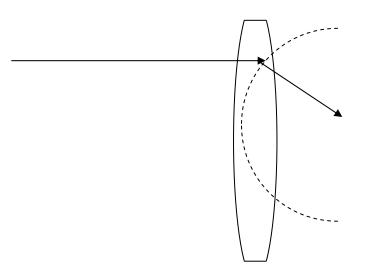
$$K\left[\left(\frac{1+M}{1-M}\right)^2 + \chi_2\right] = \frac{2}{L}$$

$$x_{\sharp} = -1 - \frac{2}{LM^3}$$

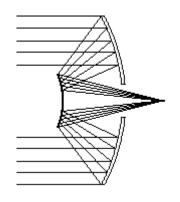
$$\chi_2 = -\left(\frac{1+M}{1-M}\right)^2 + \frac{2(1+ML)}{L(1-M)^3}$$

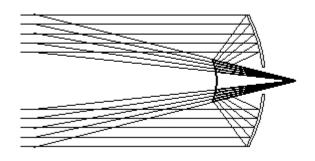


## Principal surface



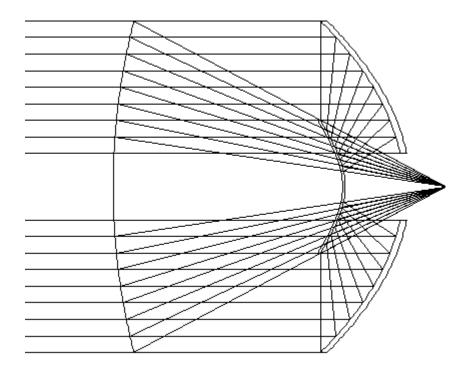
In an aplanat working at m=0
the equivalent
refracting surface
is a hemisphere







# Cassegrain's principal surface



Since the equivalent refracting surface in a Cassegrain telescope is a paraboloid then the coma of that Cassegrain is the same of a paraboloid mirror with the same focal length.

Dall-Kirkham: Spherical secondary, elliptical primary

Pressman-Camichel: Spherical primary, spheroid secondary

Consider Gregorian possible choices.



### Two mirror afocal system

#### Application to a two mirror Mersenne system

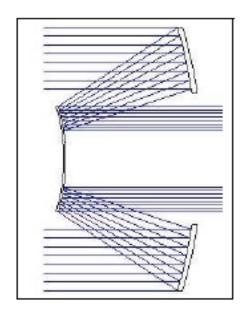
In this section we determine the aberration coefficients of a two mirror afocal system as shown in the figure We normalize the system parameters and set  $\mathcal{K} = 1$ ,  $\Phi_1 = 1$ ,  $y_1 = 1$ ,  $\overline{y_1} = 0$  and set the magnification to be m and therefore  $y_2 = m$ . We have that  $\overline{y_2} = 1 - m$ ,  $\Phi_2 = -1/m$  and therefore we can write for the conjugate factors and stop shifting parameters,

$$Y_1 = 1$$

$$Y_2 = -1$$

$$\overline{S}_1 = 0$$

$$\overline{S}_2 = \frac{y_2 \overline{y}_2 \Phi_2}{2 \mathcal{K}} = \frac{m-1}{2}$$



Two mirror afocal system

Using the formulas in the table the structural coefficients of each mirror are calculated as:

Structural aberration coefficients			
	Mirror 1	Mirror 2	
$\sigma_I$	$1+\alpha_1$	$1+\alpha_2$	
$\sigma_{I\!\!I}$	-1	$\frac{m+1}{2} + \frac{m-1}{2}\alpha_2$	
$\sigma_{I\!\!I\!I}$	1	$\left(\frac{m+1}{2}\right)^2 + \left(\frac{m-1}{2}\right)^2 \alpha_2$	
$\sigma_{I\!\!V}$	-1	-1	
$\sigma_{v}$	0	$\frac{m-1}{2}\frac{m+1}{2}\frac{m+3}{2} + \left(\frac{m-1}{2}\right)^3 \alpha_2$	



Finally the Seidel sums for the two mirror afocal system are given by:

# Seidel sums for two mirror afocal system $S_{I} = \frac{1}{4}\sigma_{I1} + \frac{1}{4}m^{4}\left(-\frac{1}{m}\right)^{3}\sigma_{I2} = \frac{1}{4}\left(\left(1 + \alpha_{1}\right) - m\left(1 + \alpha_{2}\right)\right)$ $S_{II} = \frac{1}{2}\sigma_{II1} + \frac{1}{2}m^2\left(-\frac{1}{m}\right)^2\sigma_{II2} = \frac{1}{4}(m-1)(1+\alpha_2)$ $S_{\underline{m}} = \sigma_{\underline{m}1} + \left(-\frac{1}{m}\right)\sigma_{\underline{m}2} = -\frac{1}{4}\frac{\left(m-1\right)^2}{m}\left(1+\alpha_2\right)$ $S_{IV} = \sigma_{IV1} + \left(-\frac{1}{m}\right)\sigma_{IV2} = -\frac{m-1}{m}$ $S_{V} = 2\sigma_{V1} + 2\left(\frac{1}{m}\right)^{2}\sigma_{V2} = \frac{1}{4}\frac{m-1}{m^{2}}\left(8 + 6\left(m-1\right) + \left(m-1\right)^{2}\left(1 + \alpha_{2}\right)\right)$



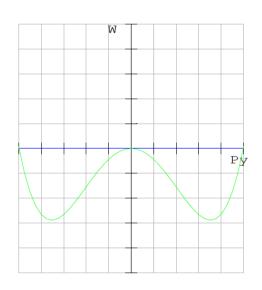
For the particular case of having the mirror as parabolic in optical shape we have that the Seidel sums simplify as:

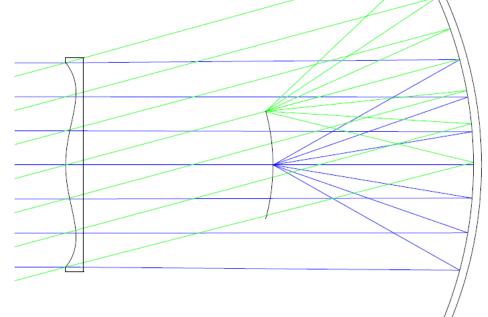
Seidel sums for afocal system using parabolas
$S_I = 0$
$S_{II} = 0$
$S_{III} = 0$
$S_{IV} = -\frac{m-1}{m}$
$S_V = \frac{1}{2} \frac{m-1}{m^2} (3m+1)$

When a system is free from spherical aberration, coma, and astigmatism it is called an anastigmatic system.



Camera Schmidt





Aspheric plate at mirror center of curvature A-bar=0 Stop aperture at aspheric plate

Note symmetry about mirror CC

No spherical aberration

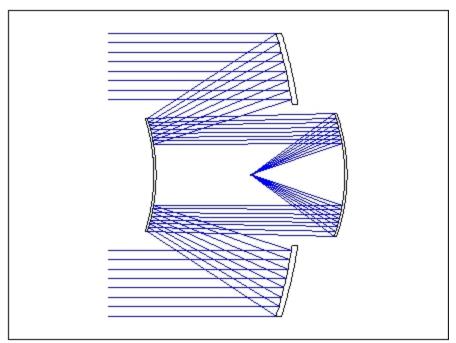
No coma

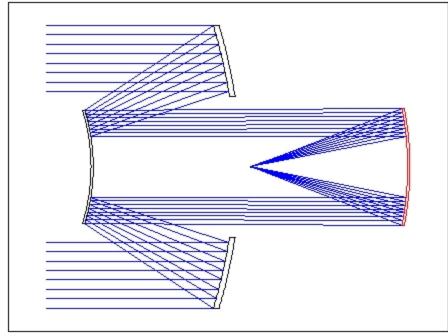
No astigmatism.

Anastigmatic over a wide field of view! Satisfies Conrady's D-d sum Note long radius on plate



# Paul-Baker system Anastigmatic-Flat field



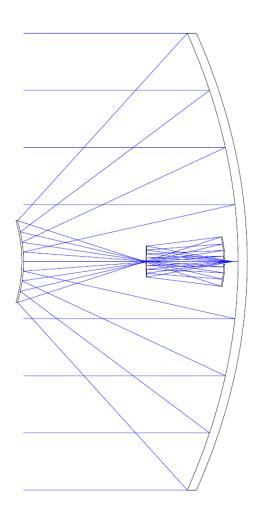


Anastigmatic
Parabolic primary
Spherical secondary and tertiary
Curved field
Tertiary CC at secondary

Anastigmatic, Flat field
Parabolic primary
Elliptical secondary
Spherical tertiary
Tertiary CC at secondary

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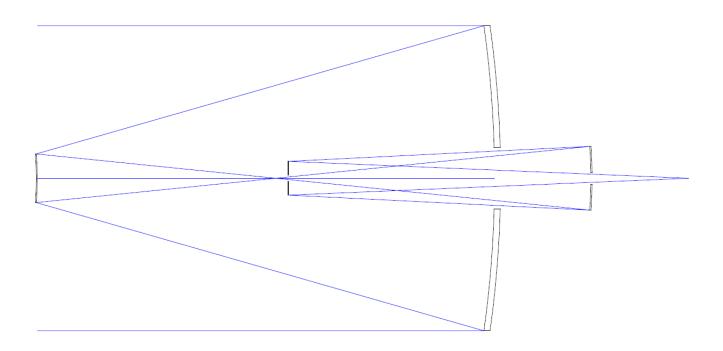
# Meinel's two stage optics concept (1985)



Large Deployable
Reflector
Second stage corrects
for errors of first stage;
fourth mirror is at the
exit pupil.



# Aplanatic, Anastigmatic, Flat-field, Orthoscopic (free from distortion, rectilinear, JS 1987)



Spherical primary telescope.

The quaternary mirror is near the exit pupil. Spherical aberration and Coma are then corrected with a single aspheric surface. The Petzval sum is zero. If more aspheric surfaces are allowed then more aberrations can be corrected.

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